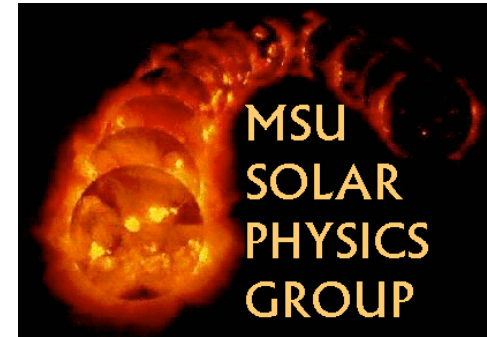


SPD June 16, 2003



Separators: Fault Lines in the Magnetic Field

Dana Longcope

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Graham Barnes
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Isaac Klapper
KD Leka

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Aad van Ballegooijen
Brian Welsch
NASA
NSF/ATM
AFOSR

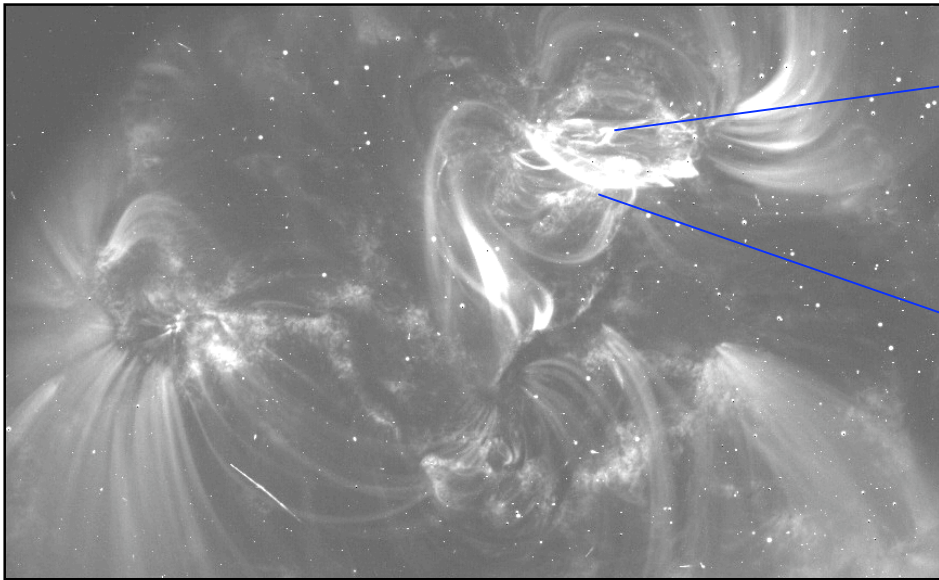
The Changing Magnetic Field

PHOTOSPHERE

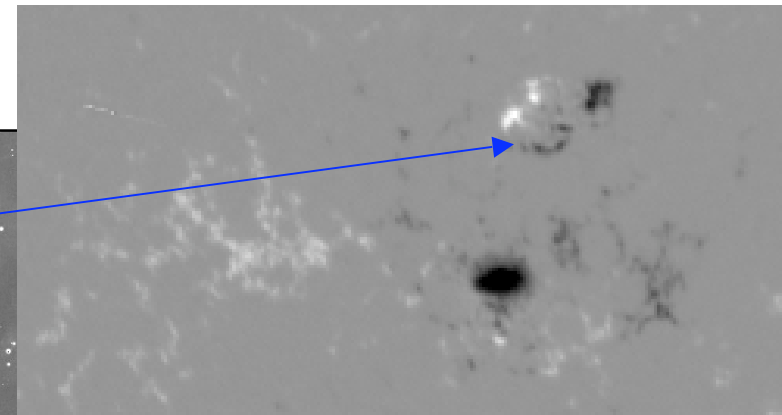
THE CORONA

SoHO MDI

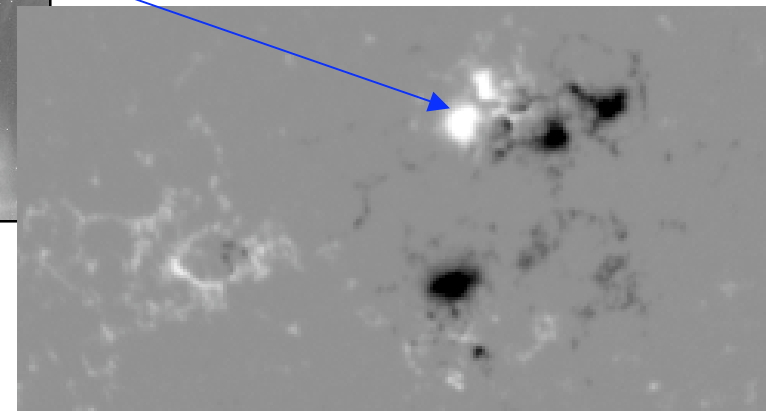
TRACE 171A



8/11/01 9:25 UT [\(movie\)](#)



8/10/01 12:51 UT

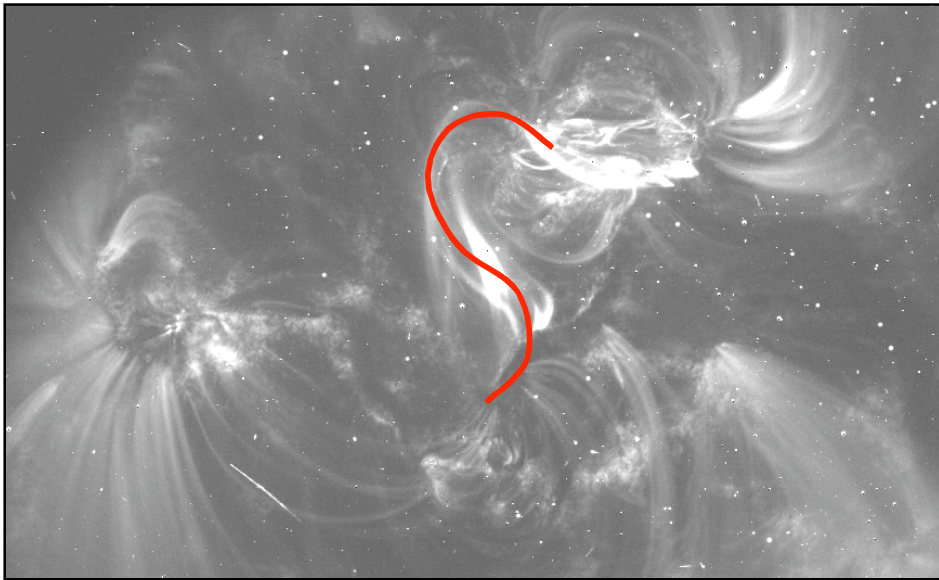


8/11/01 17:39 UT

Is this Reconnection?

THE CORONA

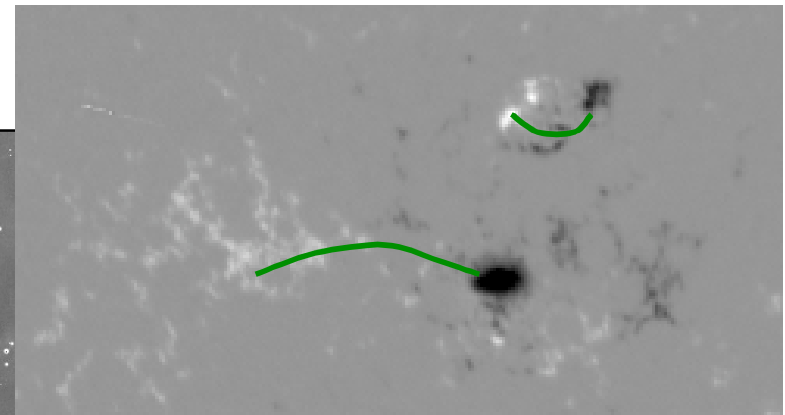
TRACE 171A



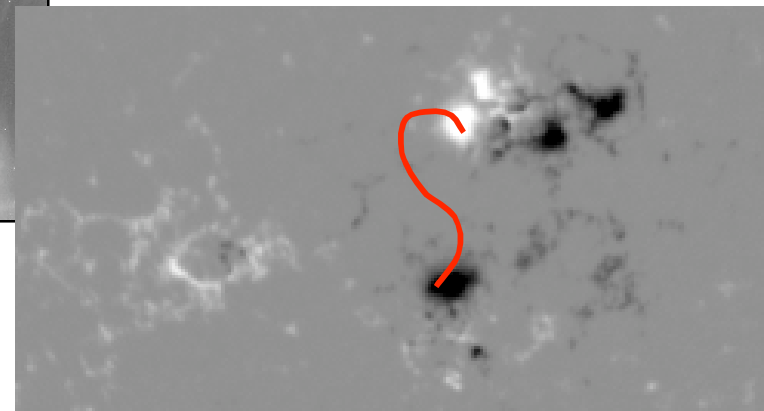
8/11/01 9:25 UT (movie)

PHOTOSPHERE

SoHO MDI



8/10/01 12:51 UT

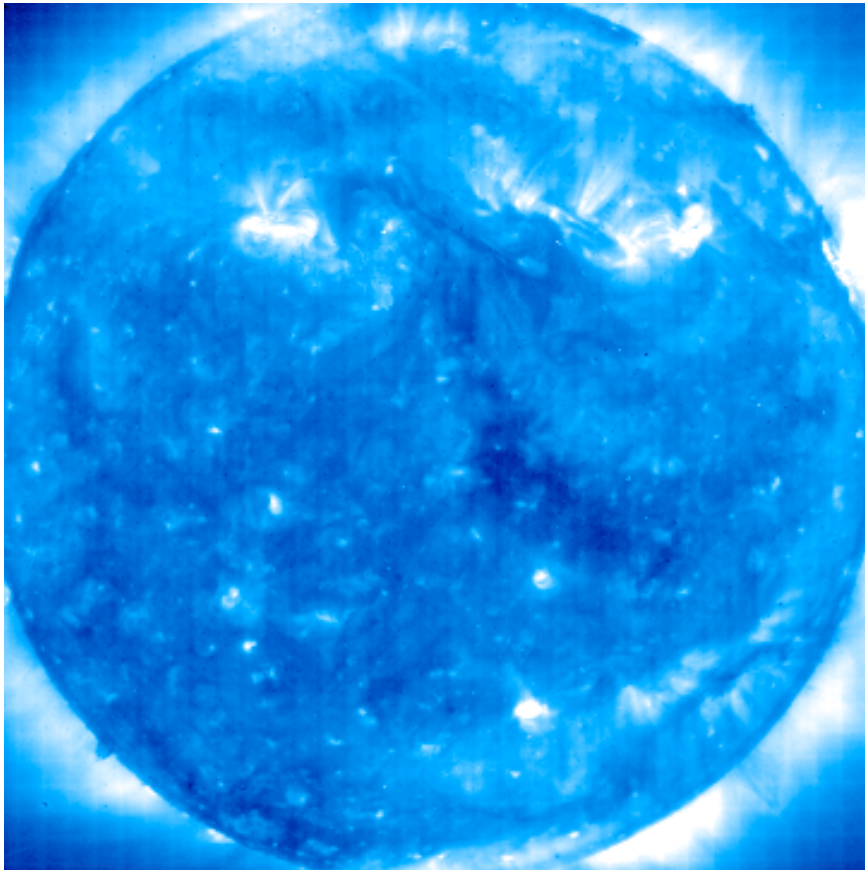


8/11/01 17:39 UT

Outline

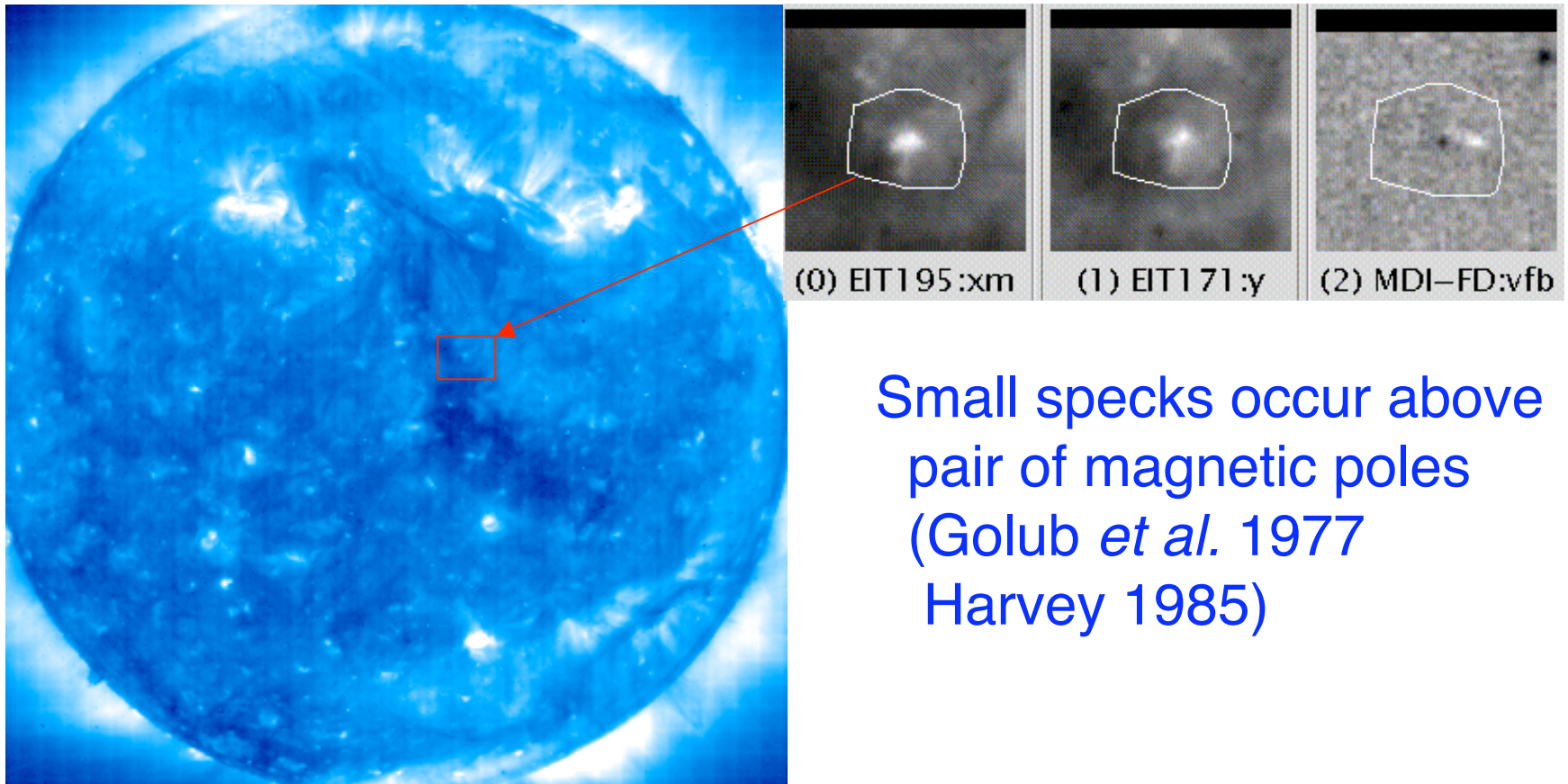
1. The XBP: A simple example of 3d reconnection
2. Quantifying Reconnection
3. Numerical simulation
4. A more complex example

Example: X-ray bright points



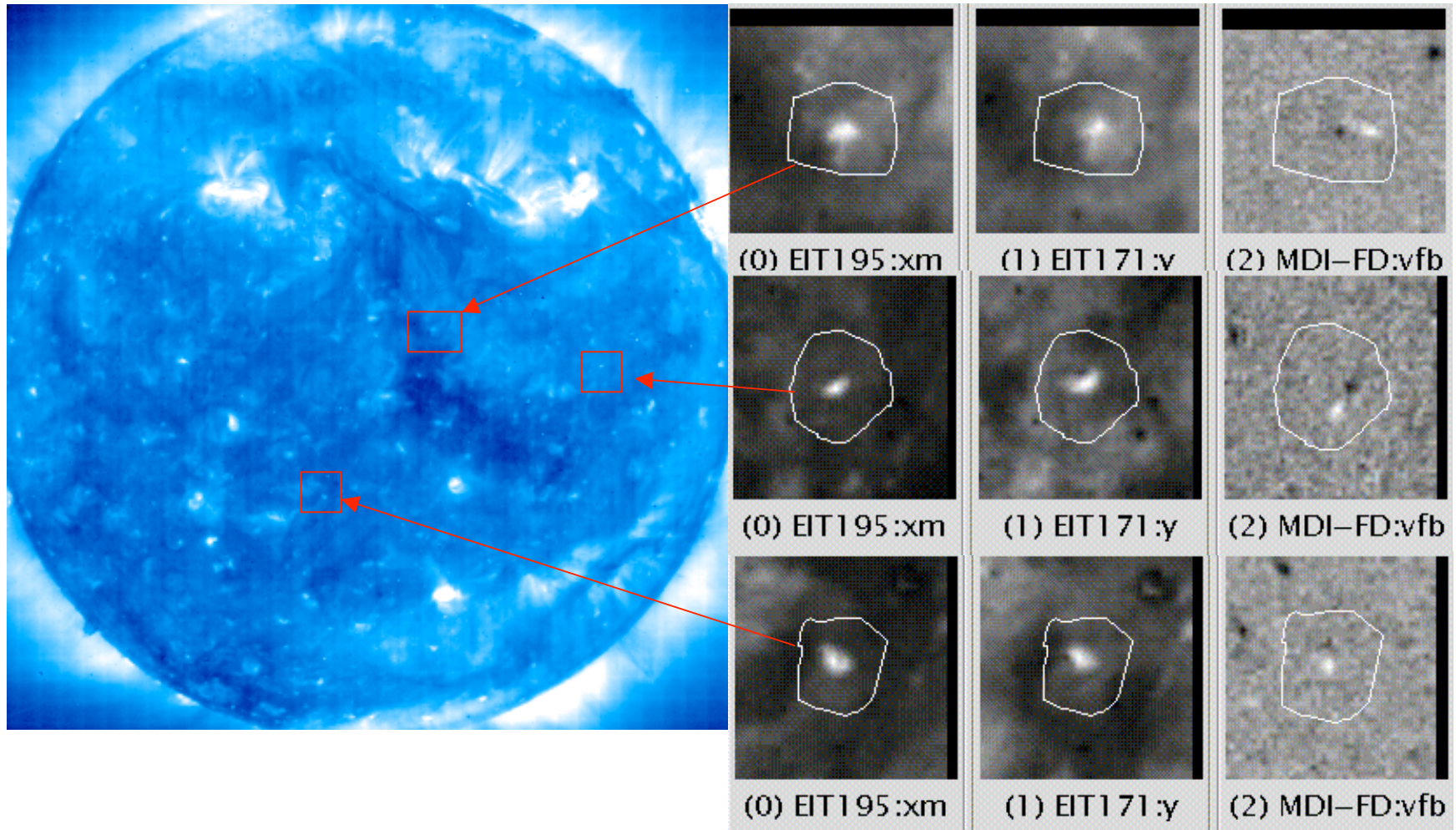
EIT 195A image of
“quiet” solar corona

Example: X-ray bright points

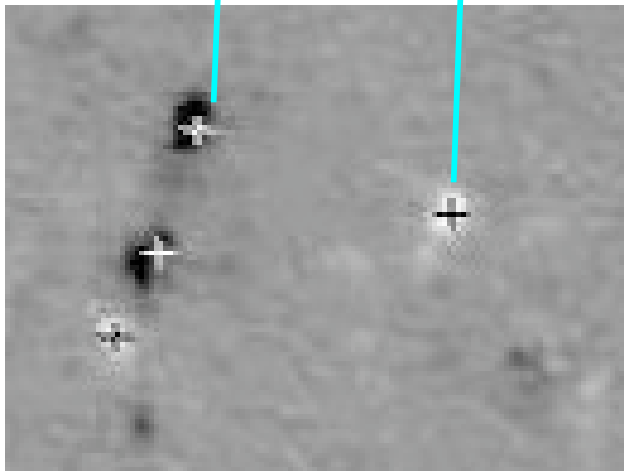
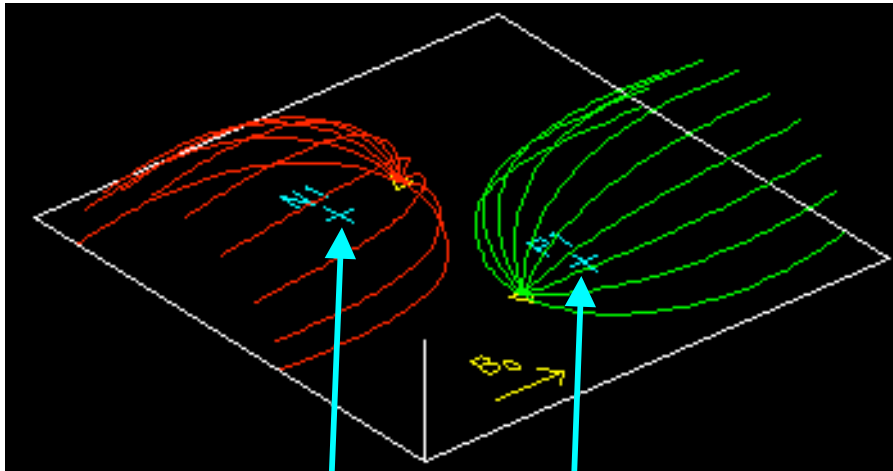


Small specks occur above pair of magnetic poles
(Golub *et al.* 1977
Harvey 1985)

Example: X-ray bright points

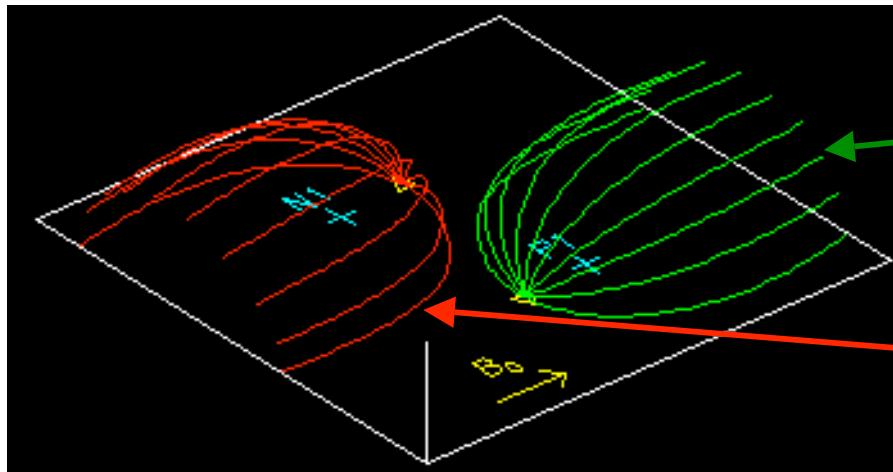


When 2 Poles Collide



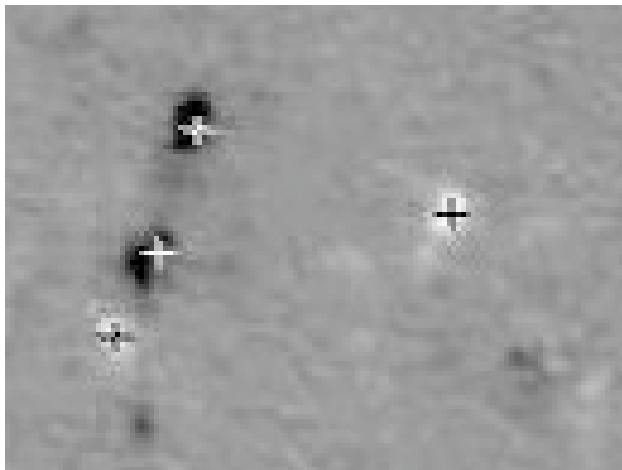
Photospheric flux
concentrations
sources of
coronal field

When 2 Poles Collide

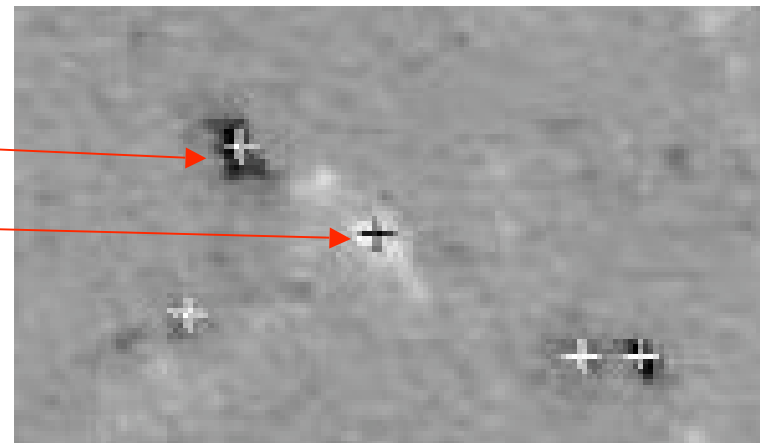
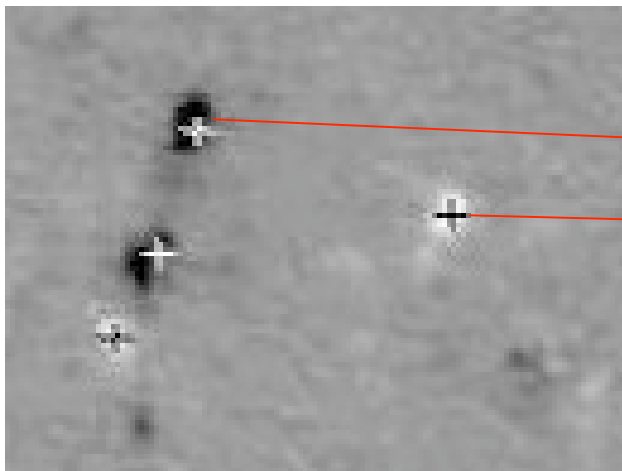
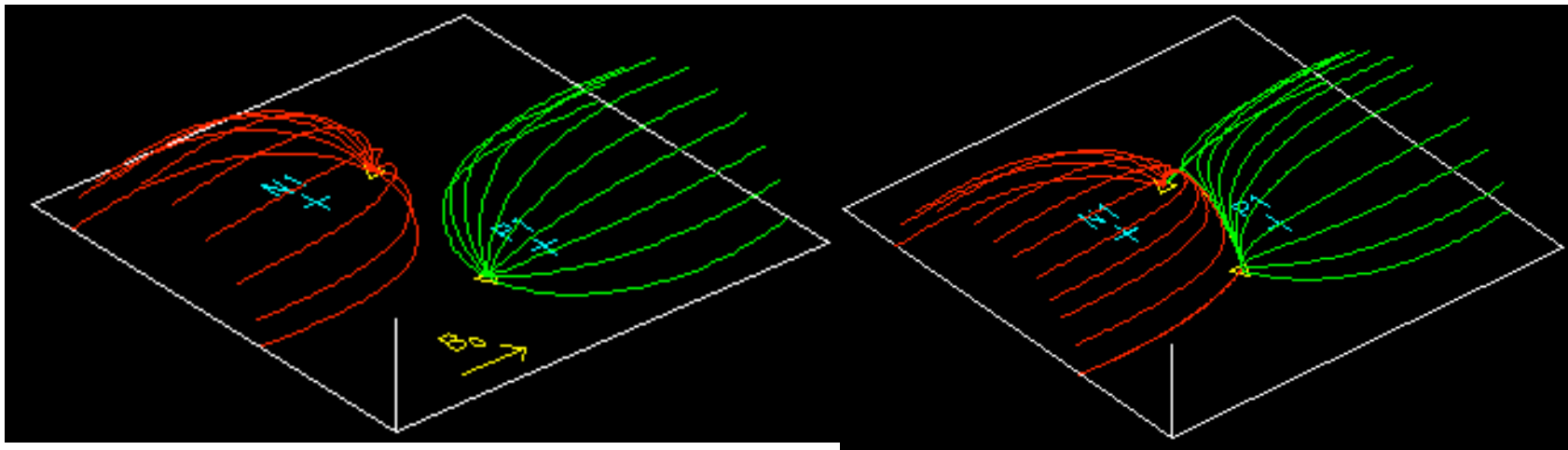


All field lines from
positive source P1

All field lines to
negative source N1

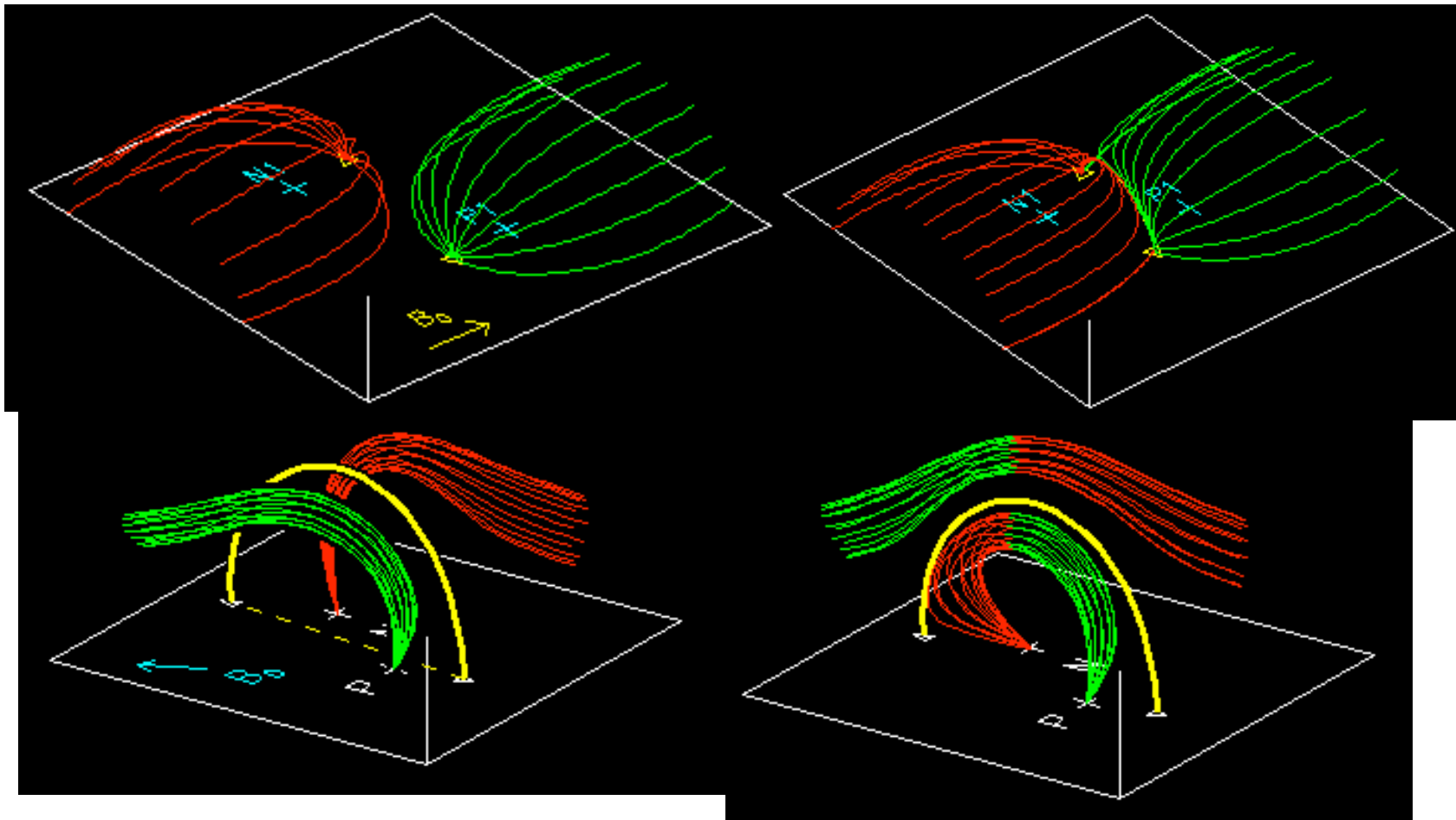


When 2 Poles Collide



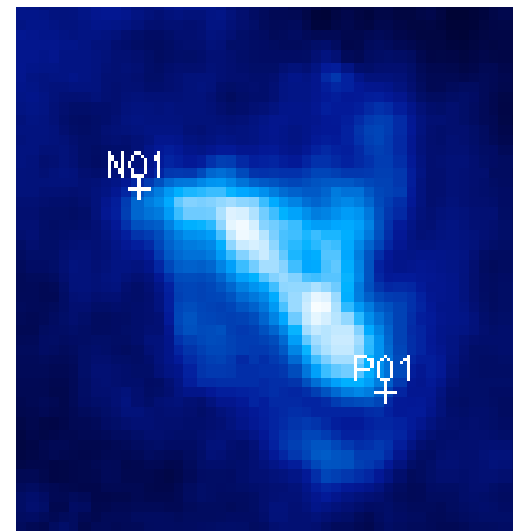
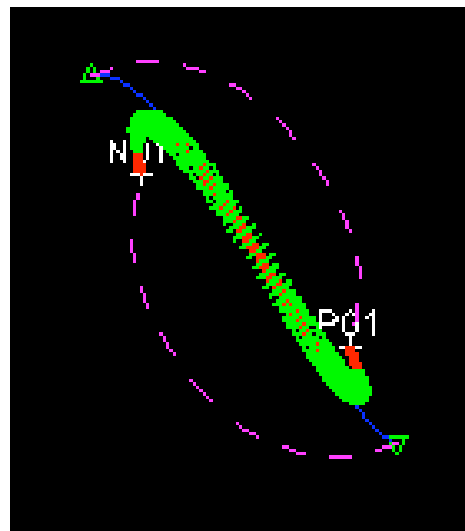
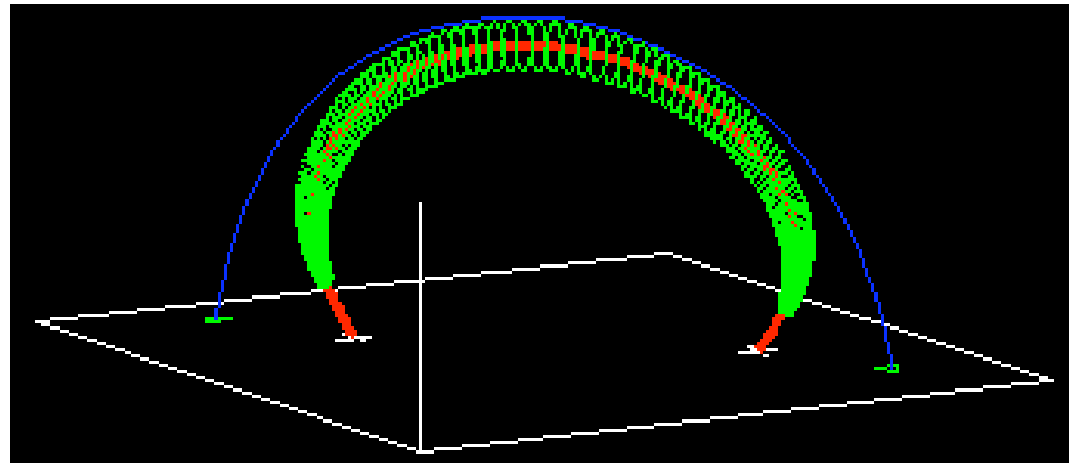
Poles approach: domains intersect

When 2 Poles Collide



Reconnection = new field lines

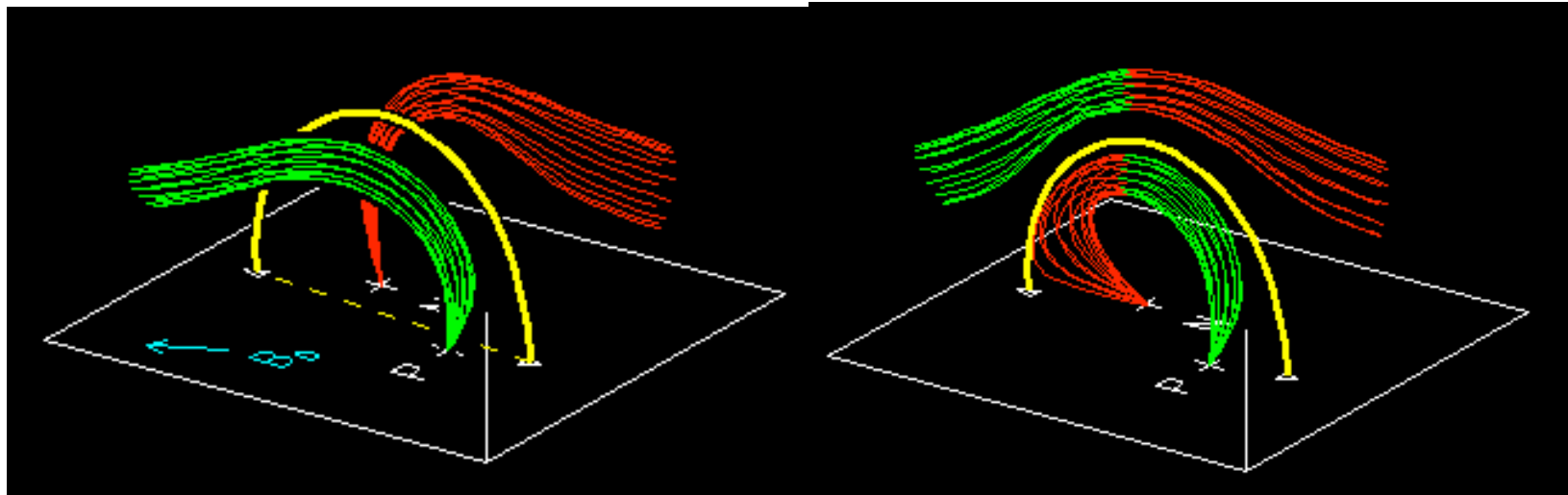
Post-reconnection Flux Tube



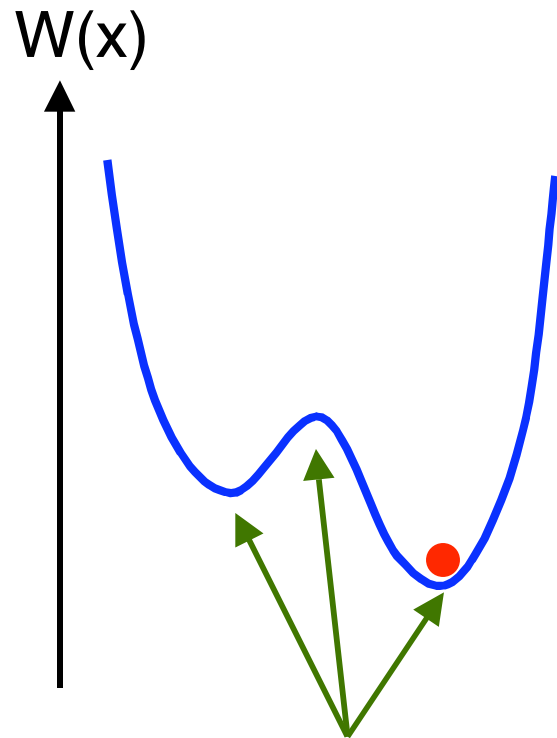
TRACE observations 6/17/98 (Kankelborg & Longcope 1999)

Quantifying Reconnection

- Why does it release energy?
- How much energy can it release?
- What about reconnection in complex magnetic fields?

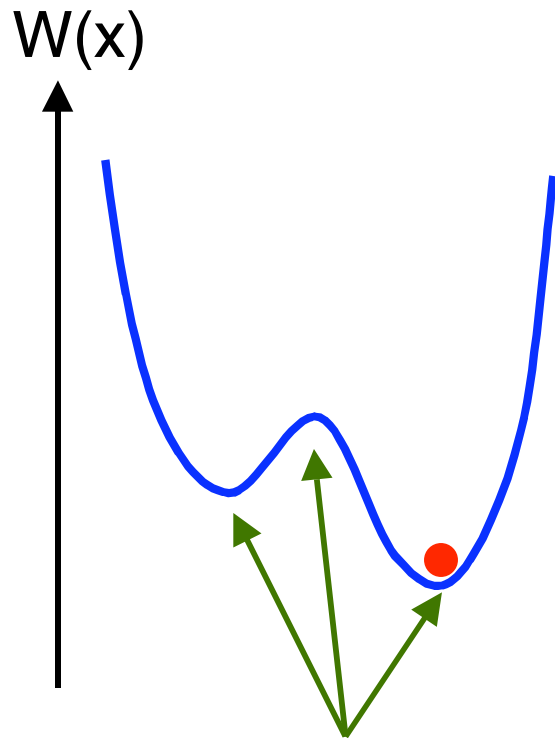


Quasi-static Evolution

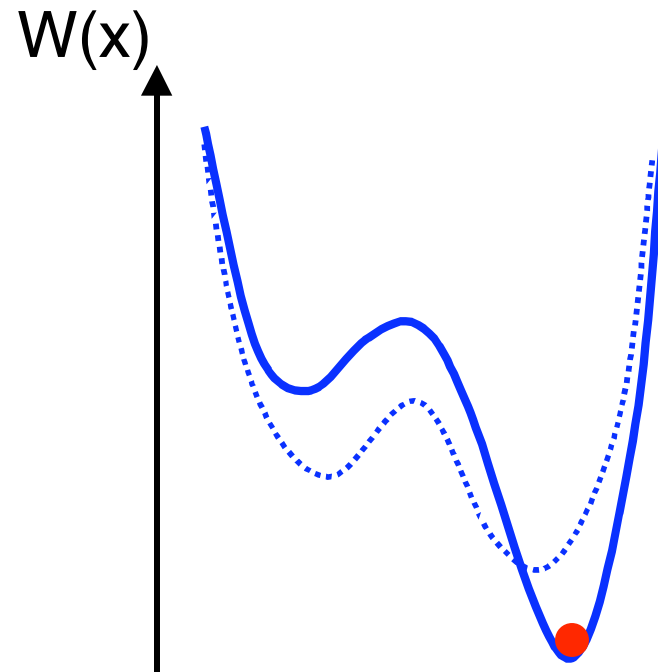


Equilibrium: $W'(x)=0$

Quasi-static Evolution



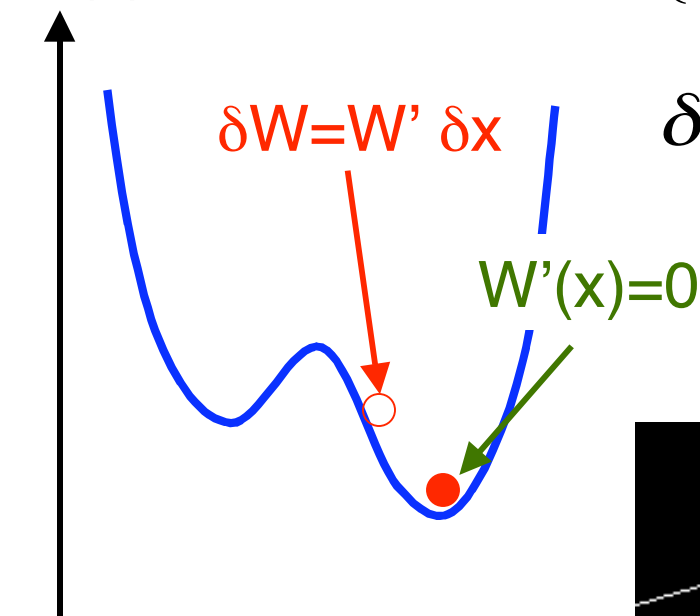
Equilibrium: $W'(x)=0$



$W(x)$ evolves ... **SLOWLY**

Equilibrium: Minimum Energy

$W(x)$

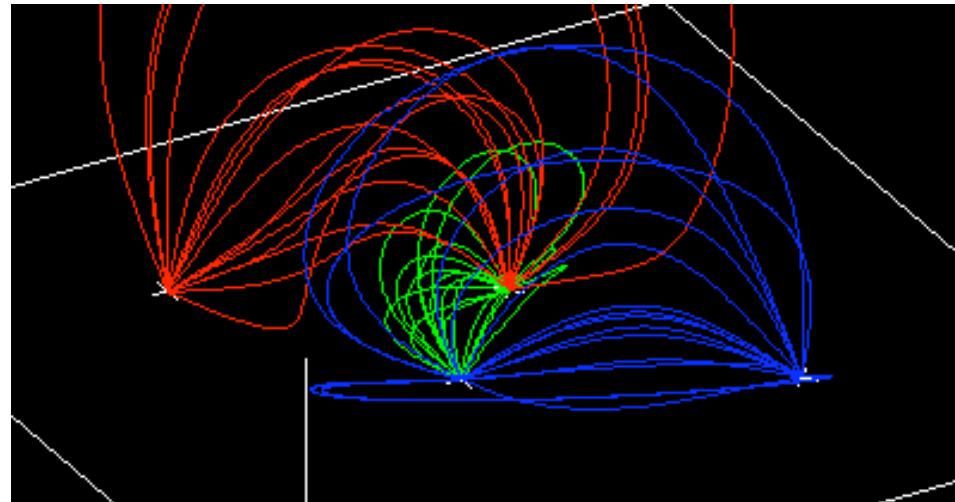


$$W\{\mathbf{A}(x)\} = \frac{1}{8\pi} \int |\nabla \times \mathbf{A}|^2 d^3x$$

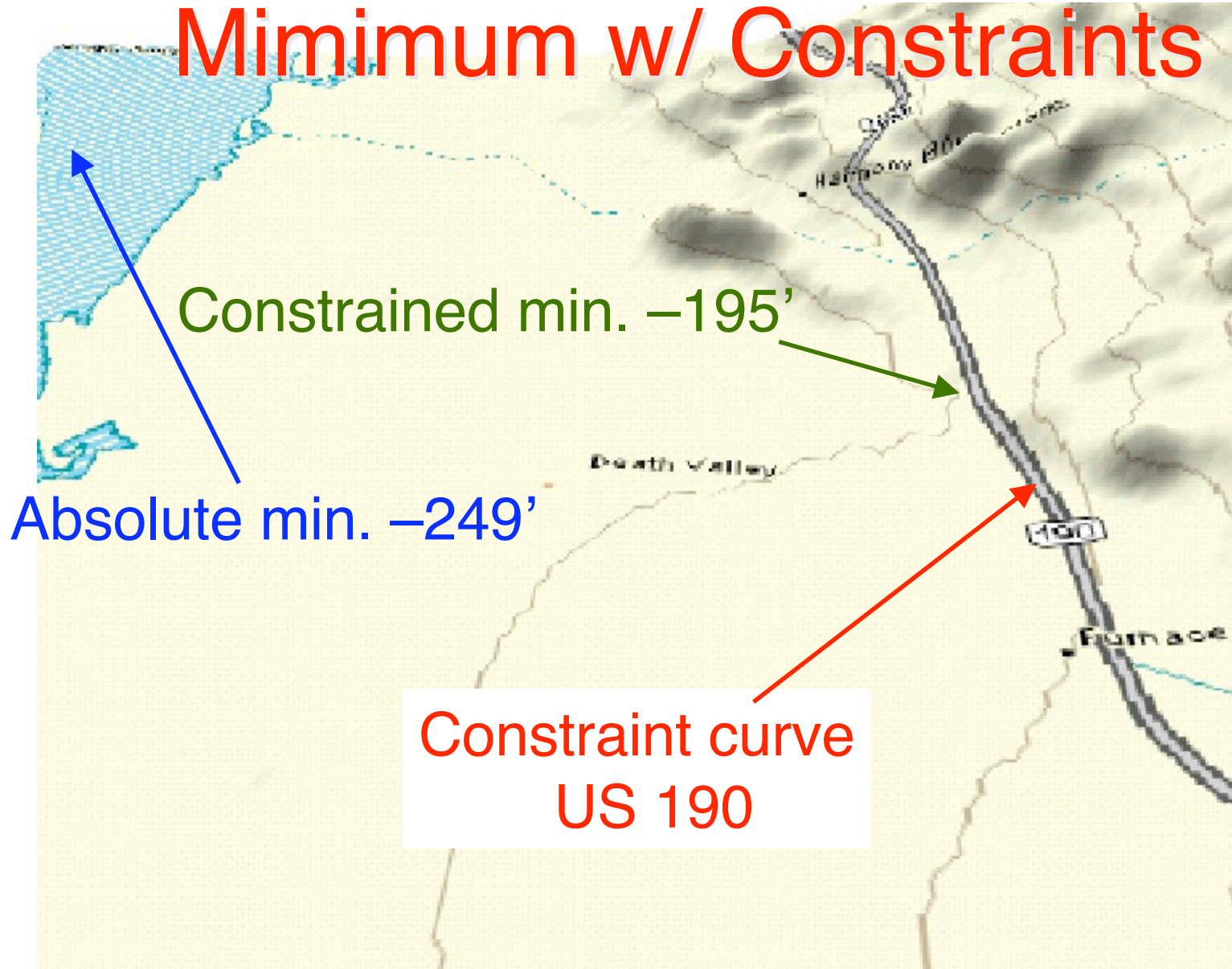
$$\delta W = \frac{1}{4\pi} \int (\nabla \times \mathbf{B}) \cdot \delta \mathbf{A} d^3x$$

$$\nabla \times \mathbf{B} = 0$$

potential



Minimum w/ Constraints



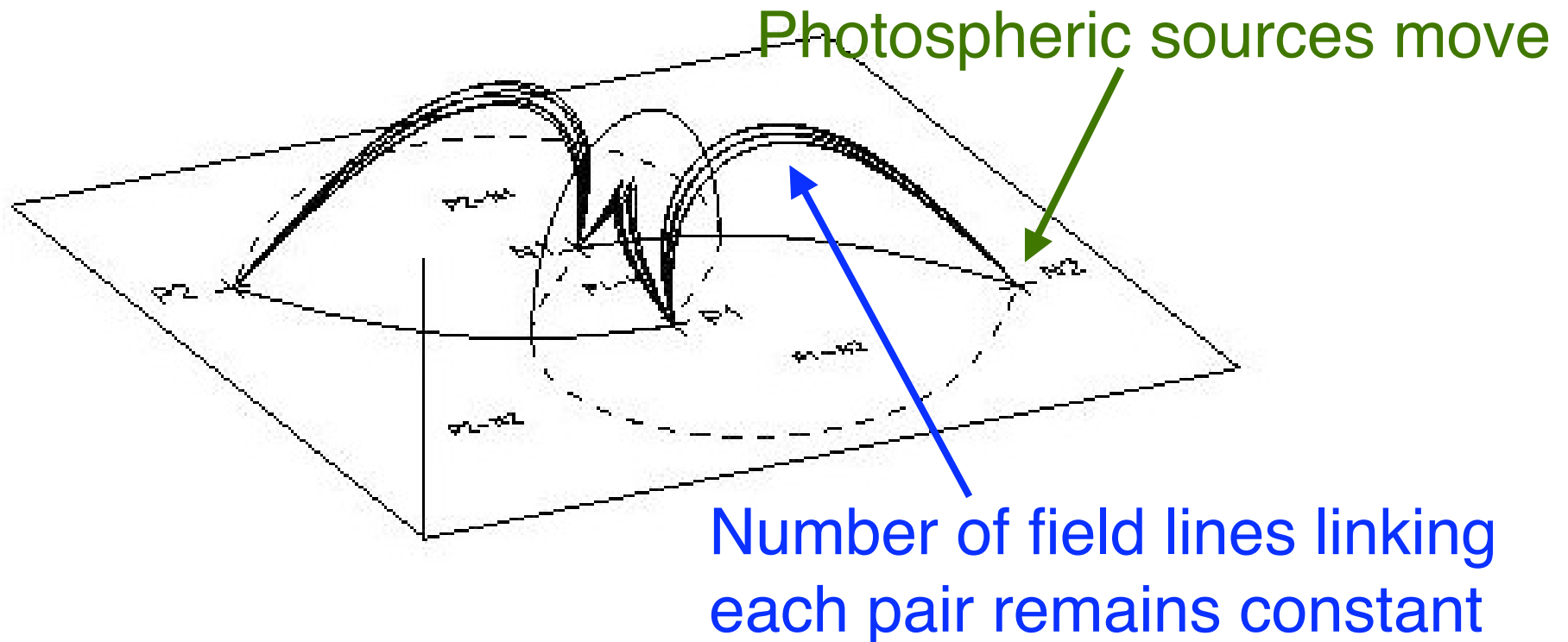
Absolute min. -249'

Constrained min. -195'

Constraint curve
US 190

A new type of constraint...

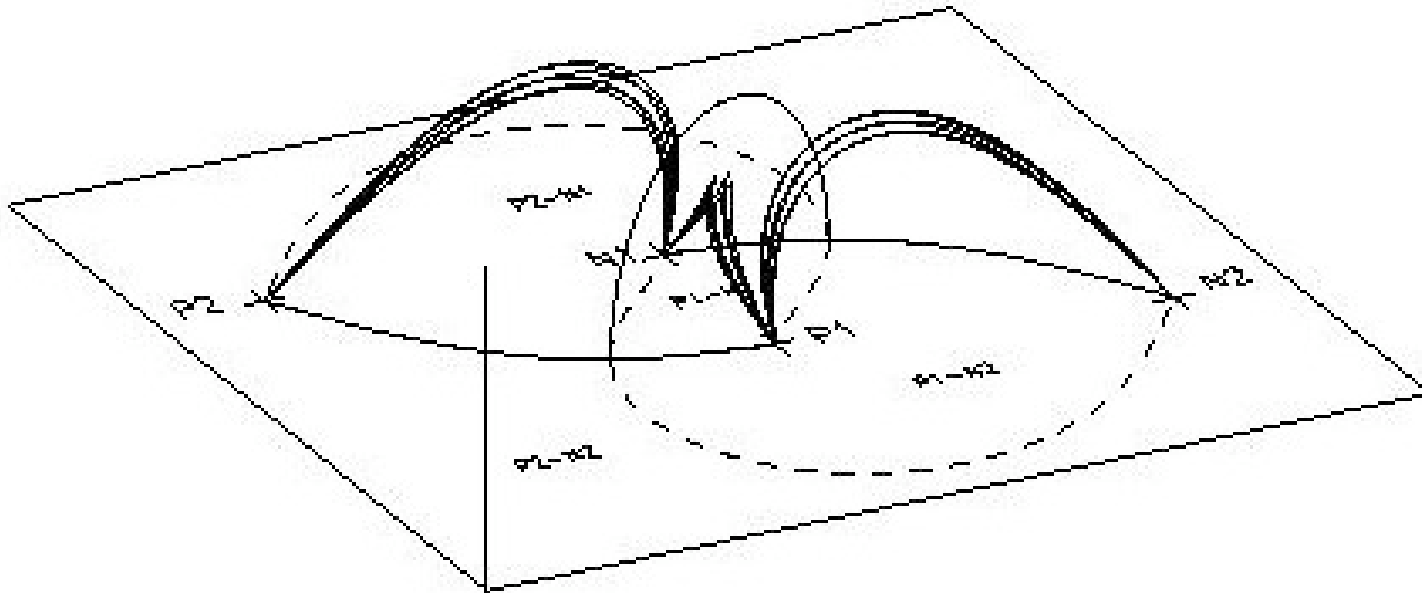
(Longcope 2001, Longcope & Klapper 2002)



No reconnection

A new type of constraint...

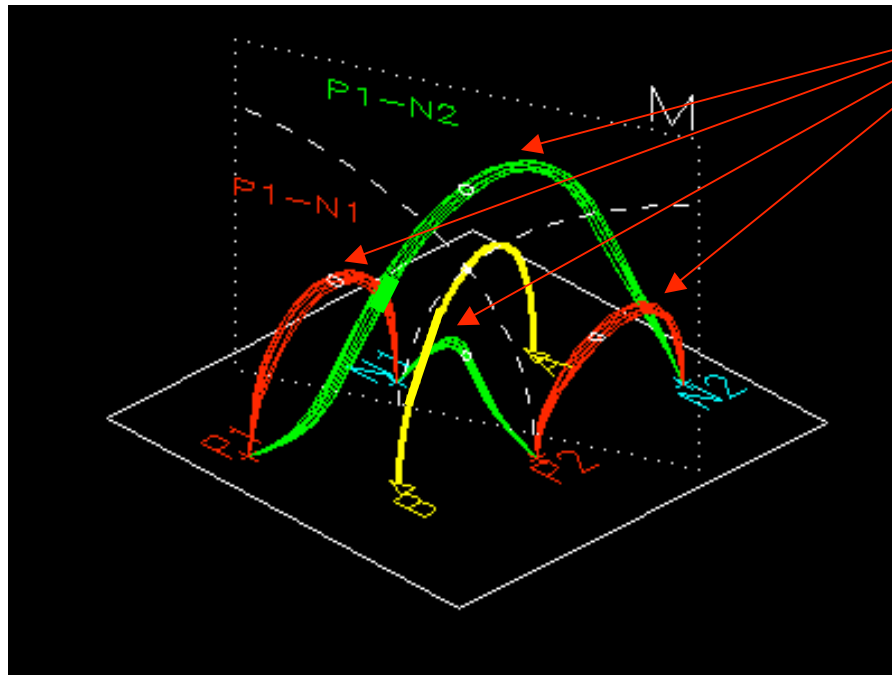
(Longcope 2001, Longcope & Klapper 2002)



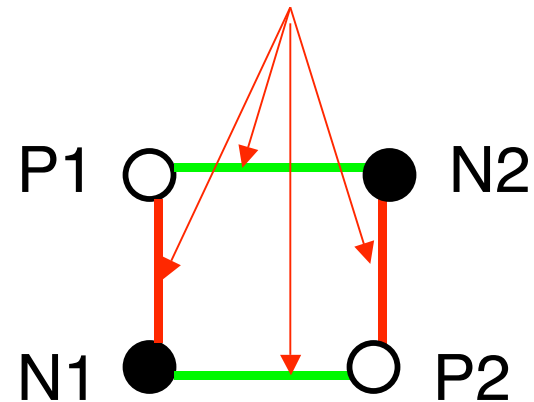
Minimize: $W\{\mathbf{A}(x)\} = \frac{1}{8\pi} \int |\nabla \times \mathbf{A}|^2 d^3x$

Subject to **flux constraints**

Separators: where domains meet

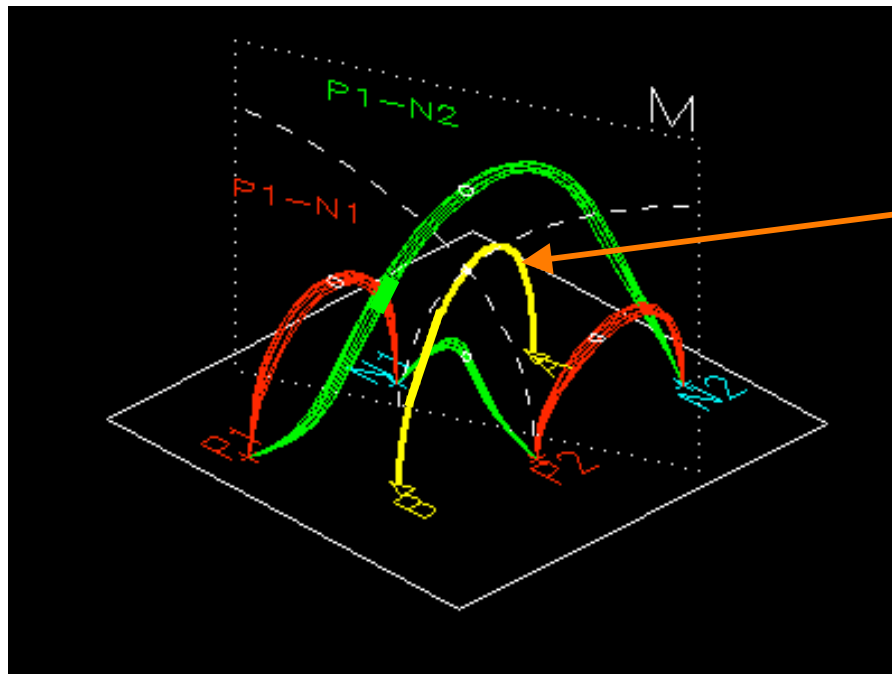


Distinct flux domains

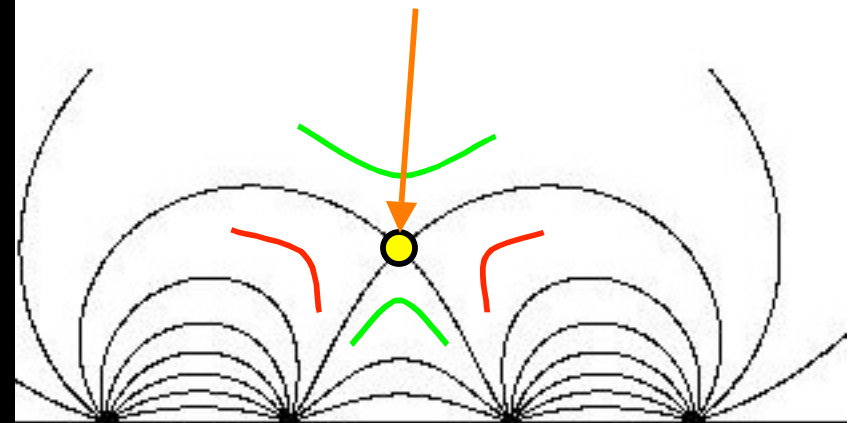


Separators: where domains meet

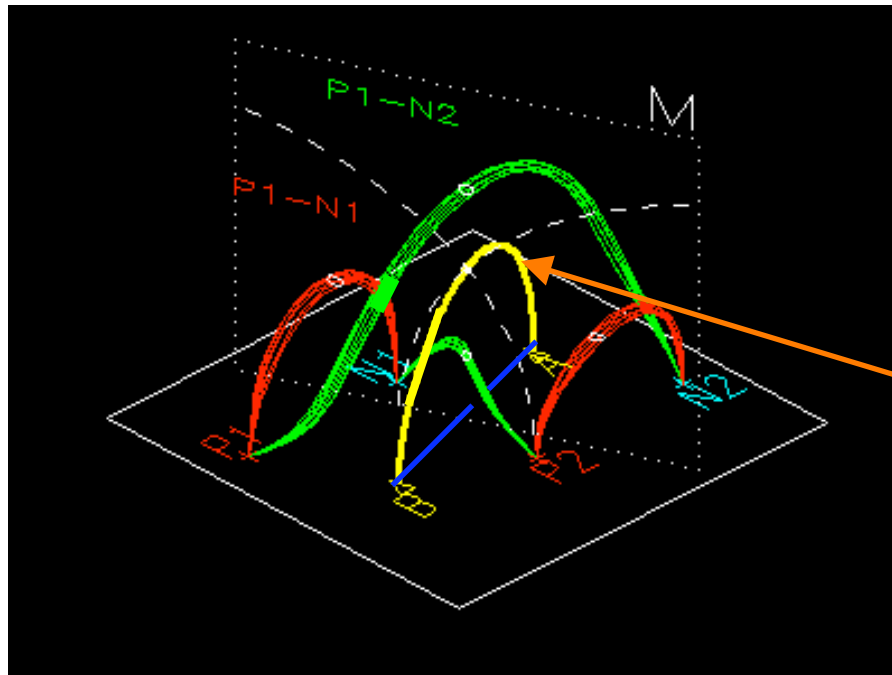
Distinct flux domains



Separator at interface



Separators: where domains meet

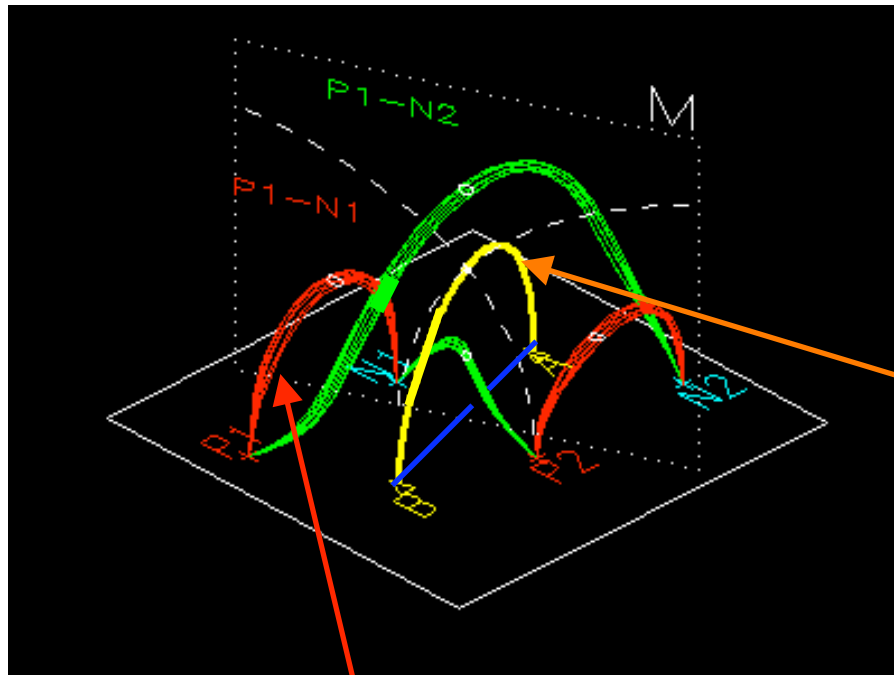


Distinct flux domains

Separator at interface

Closed loop encloses
all flux linking
P2-N1

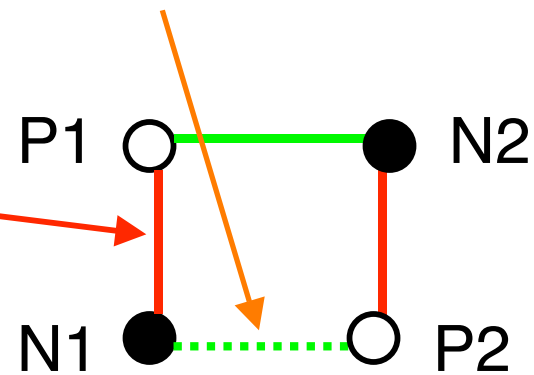
The Separator Constraint



**Constraint
only at separator**

Closed loop encloses
all flux linking
P2-N1

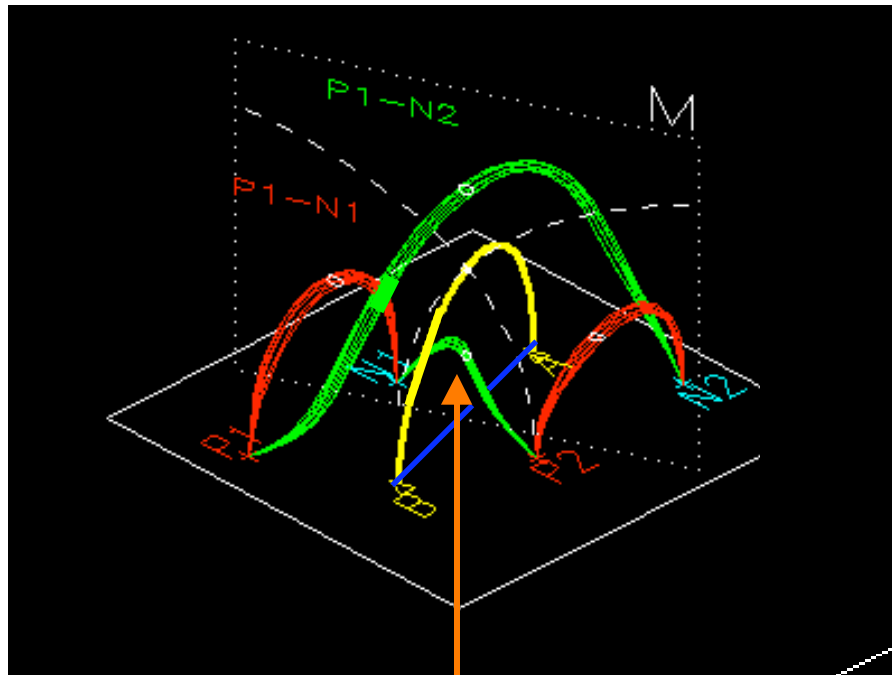
Fluxes in remaining
domains: set by BC



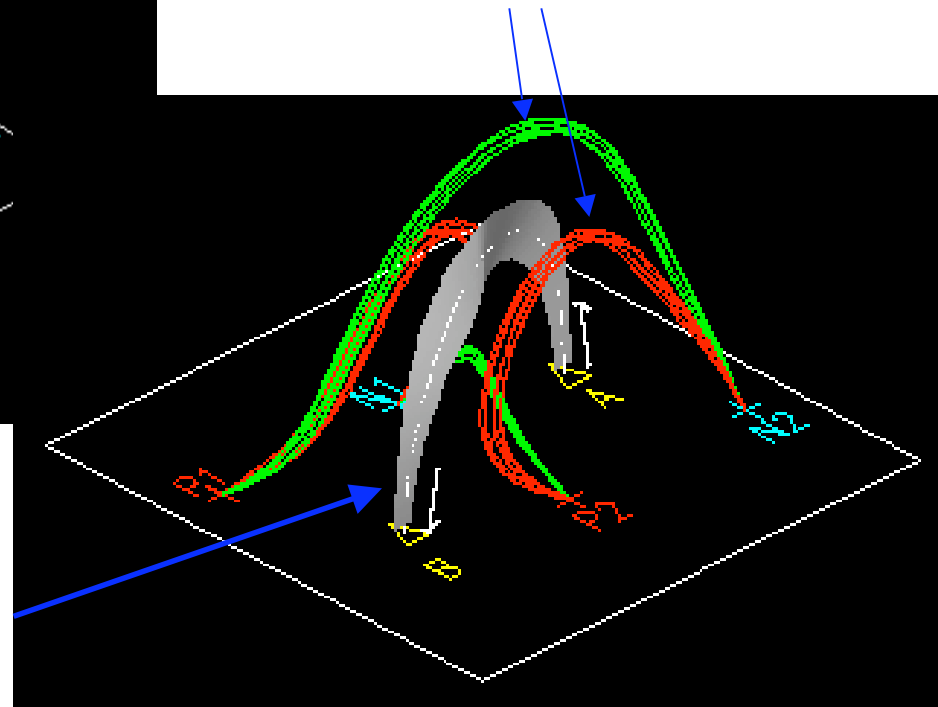
Minimum W subj. to constraint

$$\nabla \times \mathbf{B} = 0$$

Current-free within
each domain

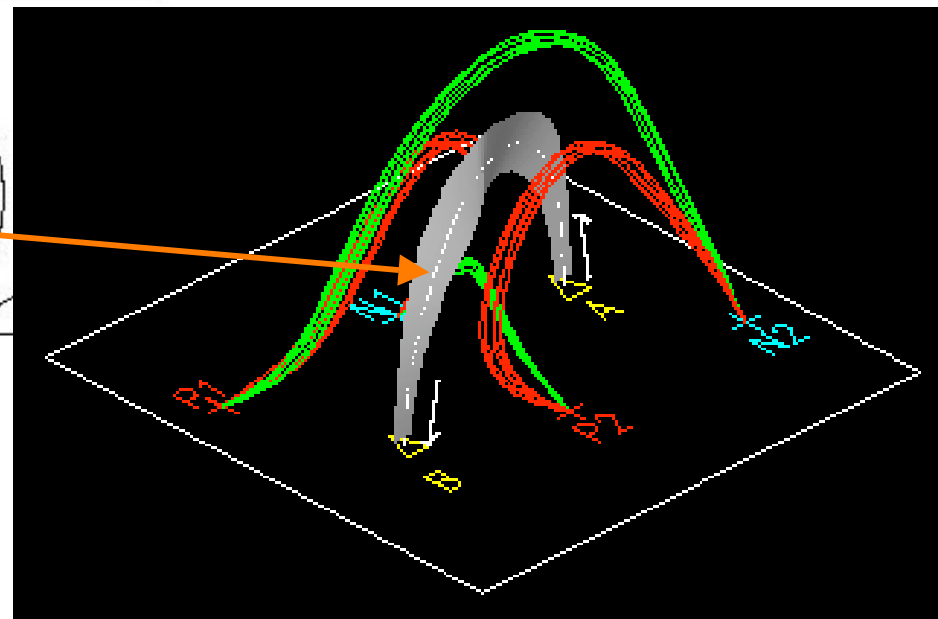
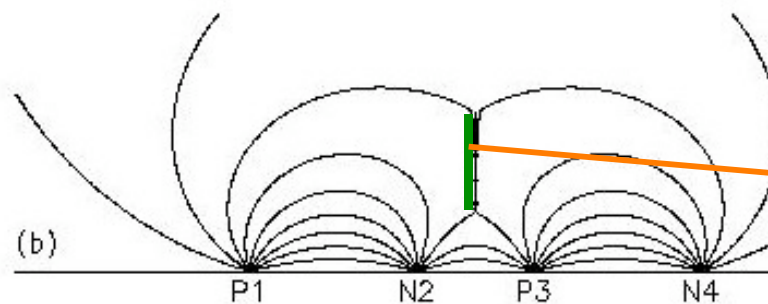
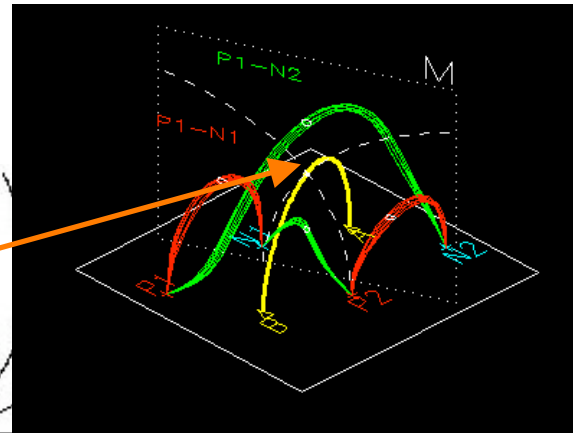
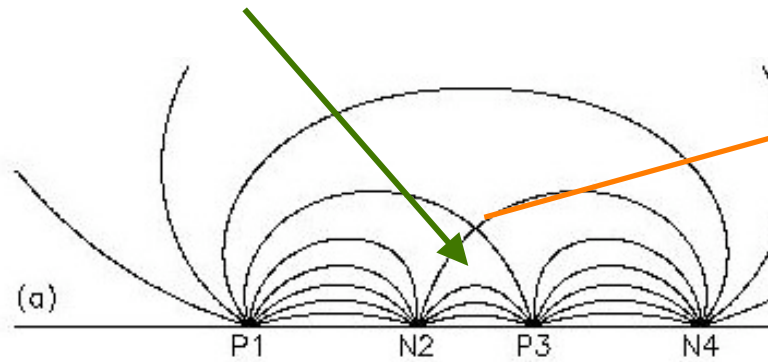


Constraint on P2-N1 flux
current sheet at separator



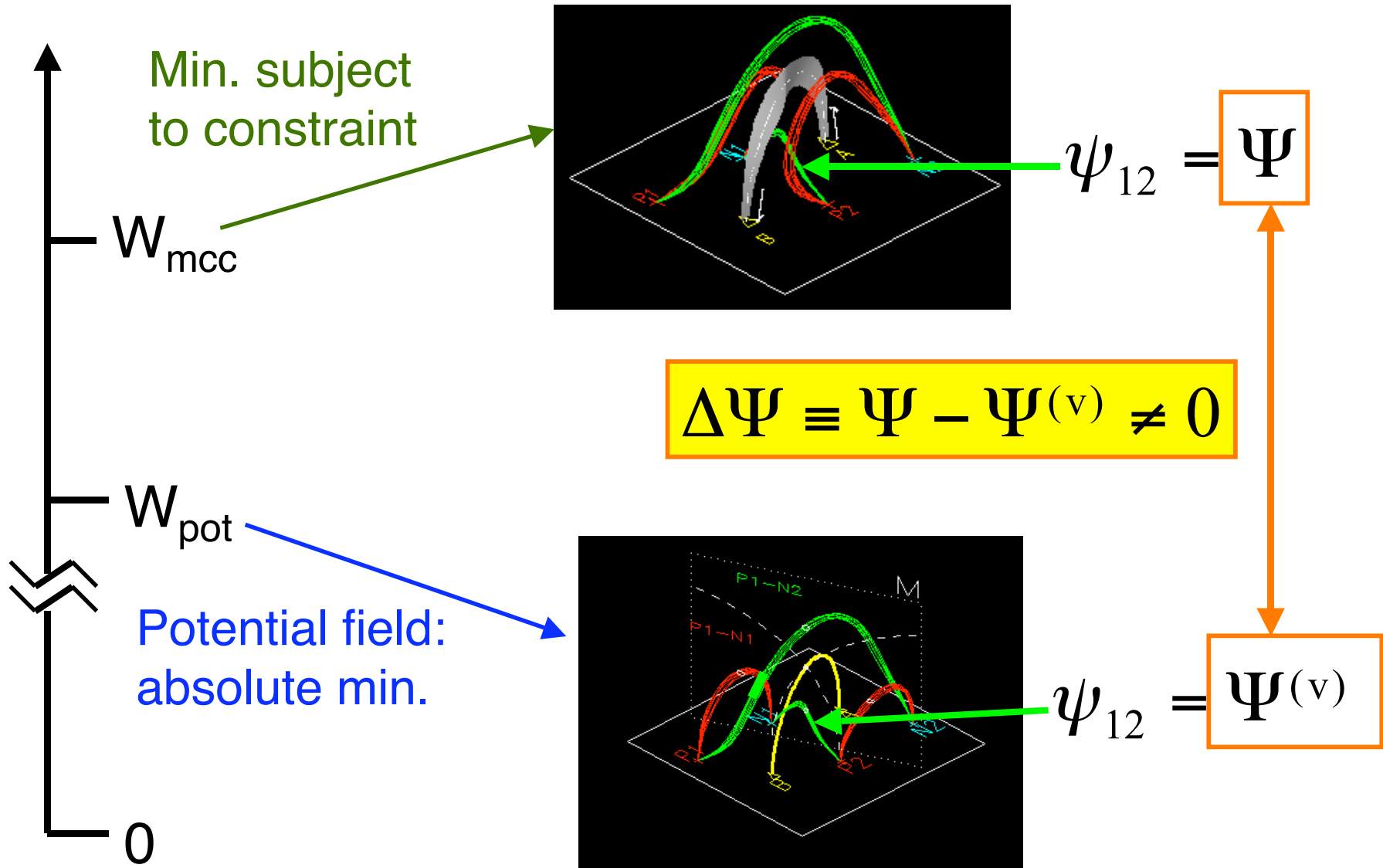
Minimum W subj. to constraint

Constraint on P3-N2 flux

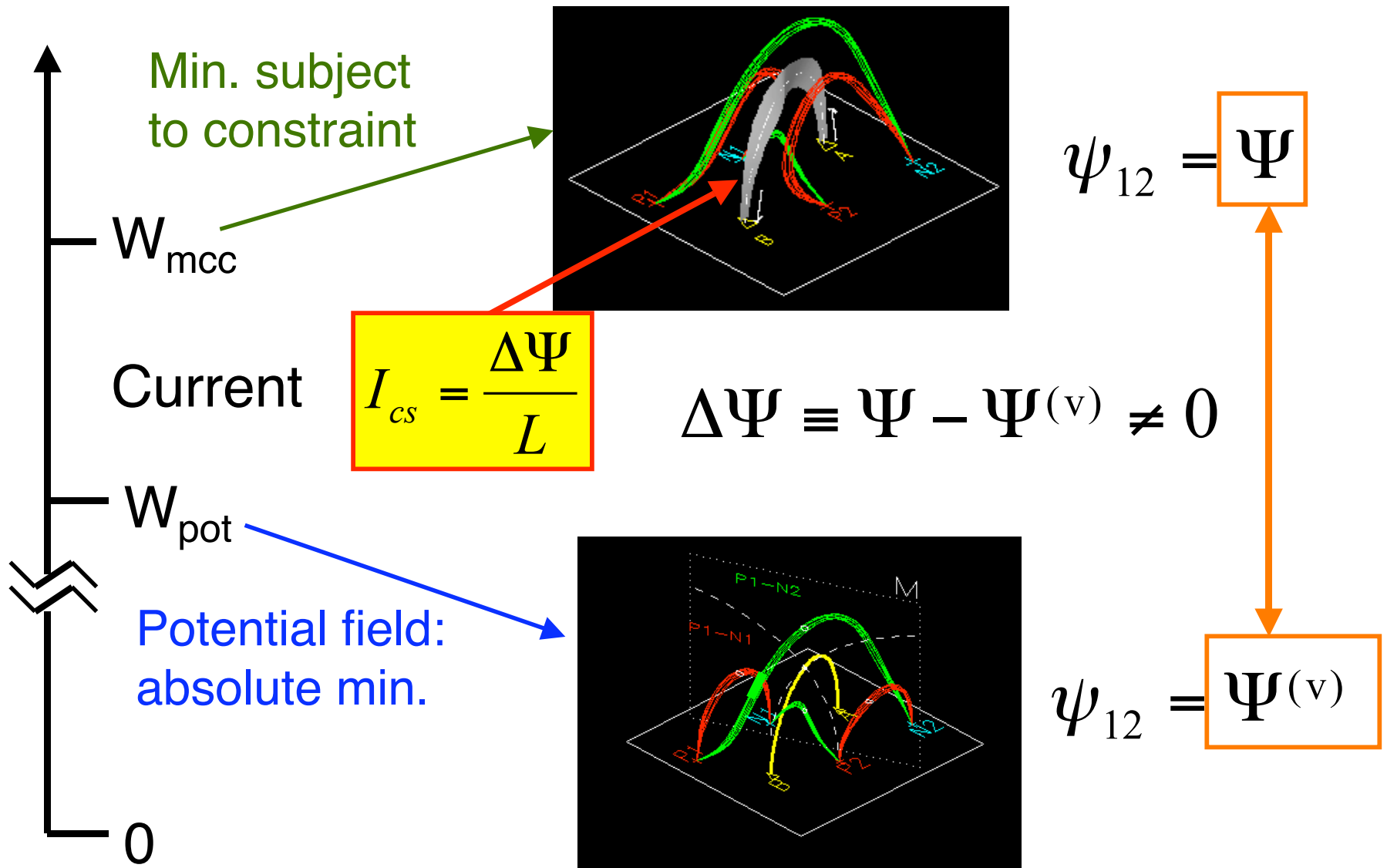


2d version: X-point
@ boundary of 4 domains
becomes current sheet

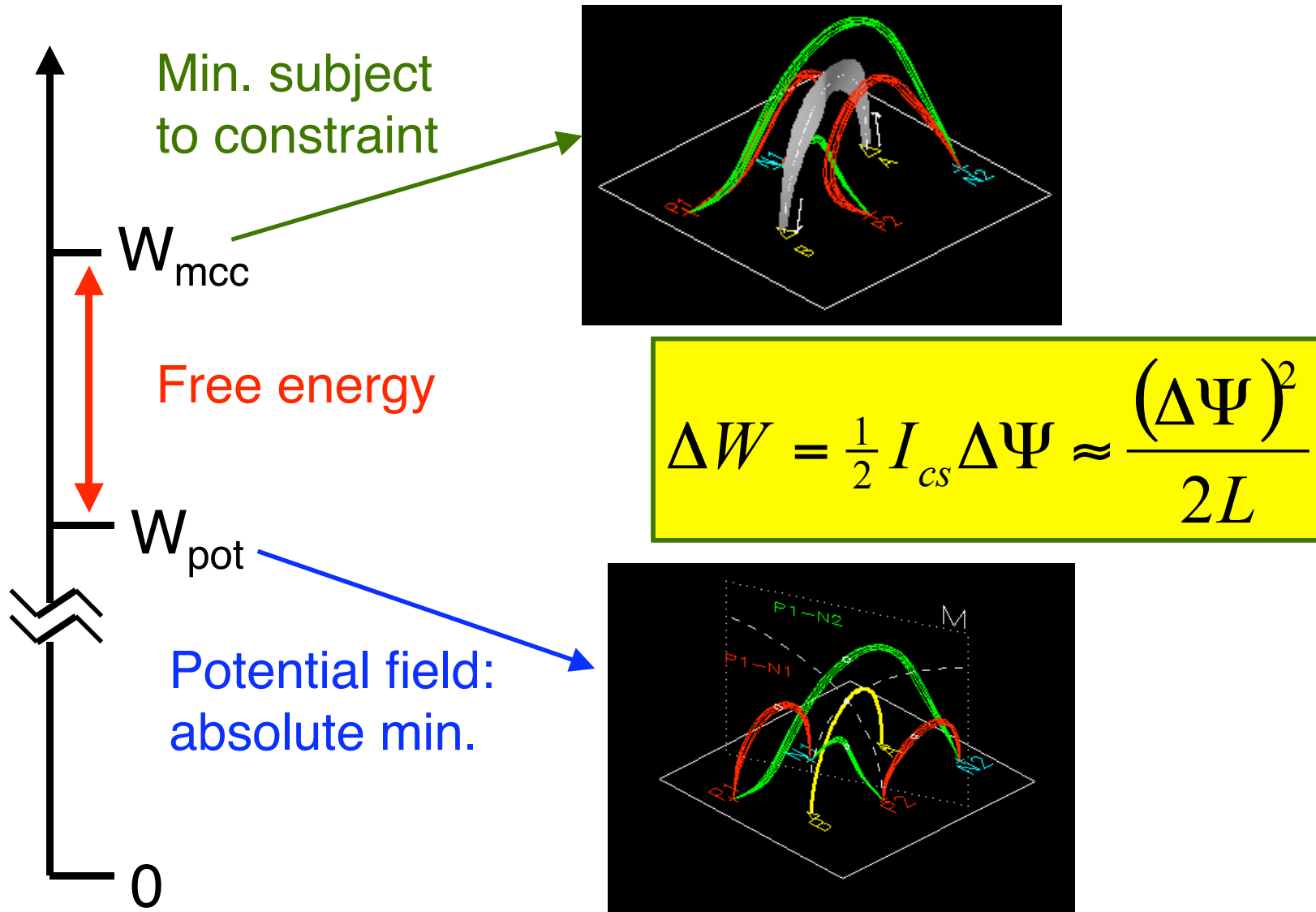
Constrained Minimum



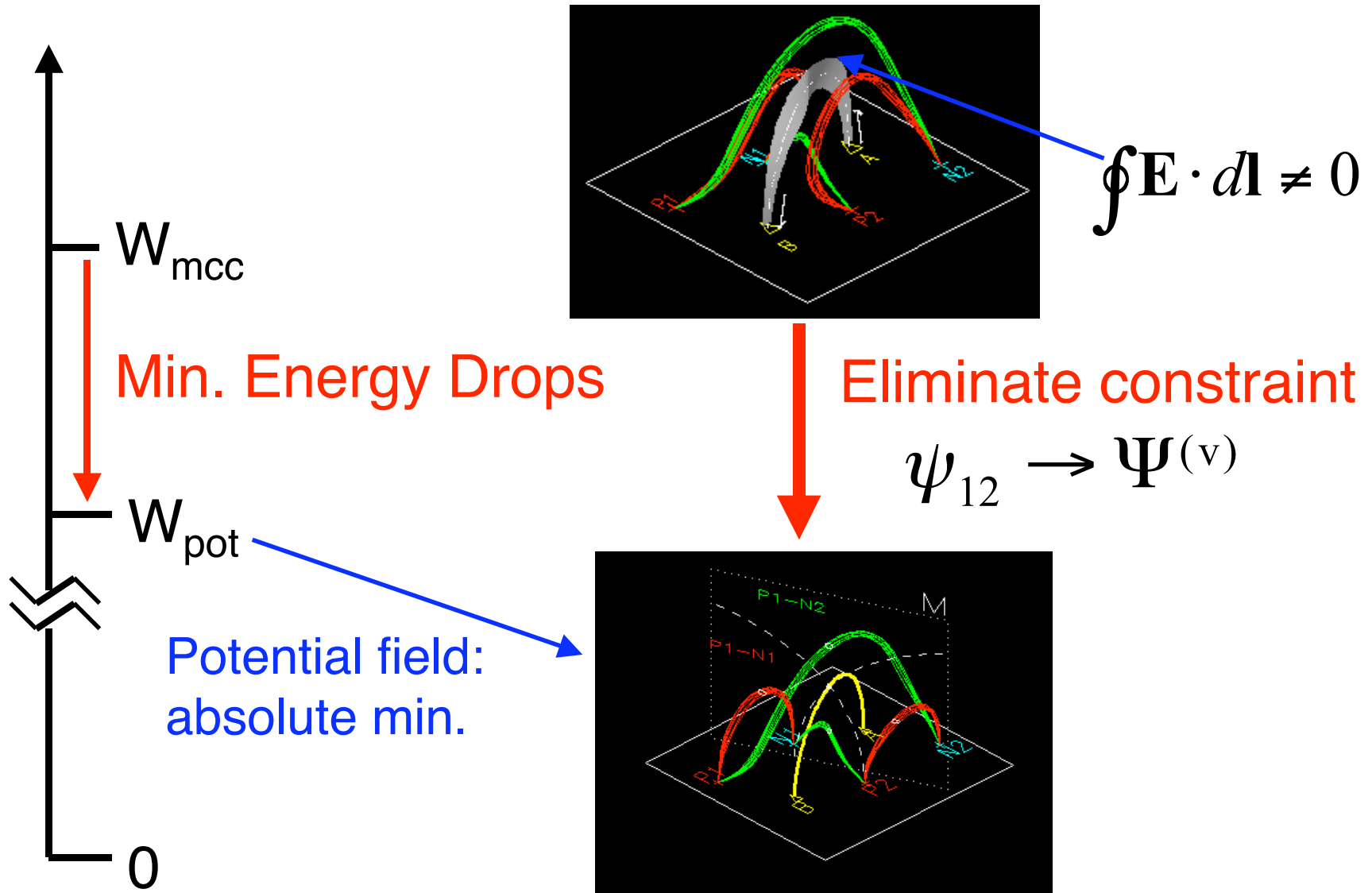
Constrained Minimum



Constrained Minimum



Reconnection

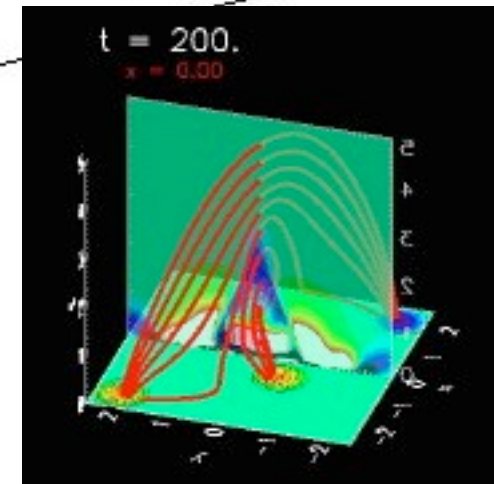
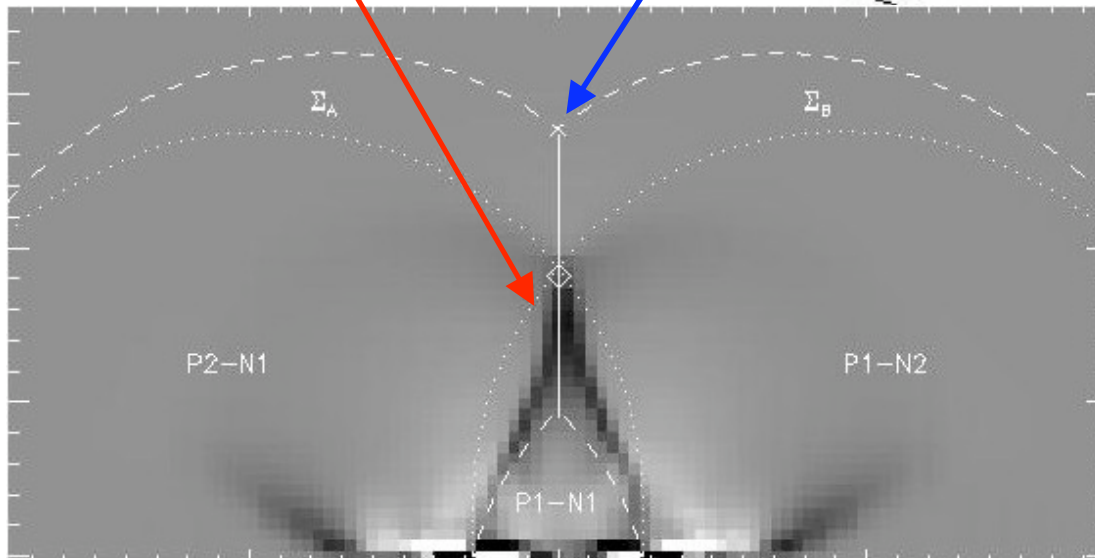
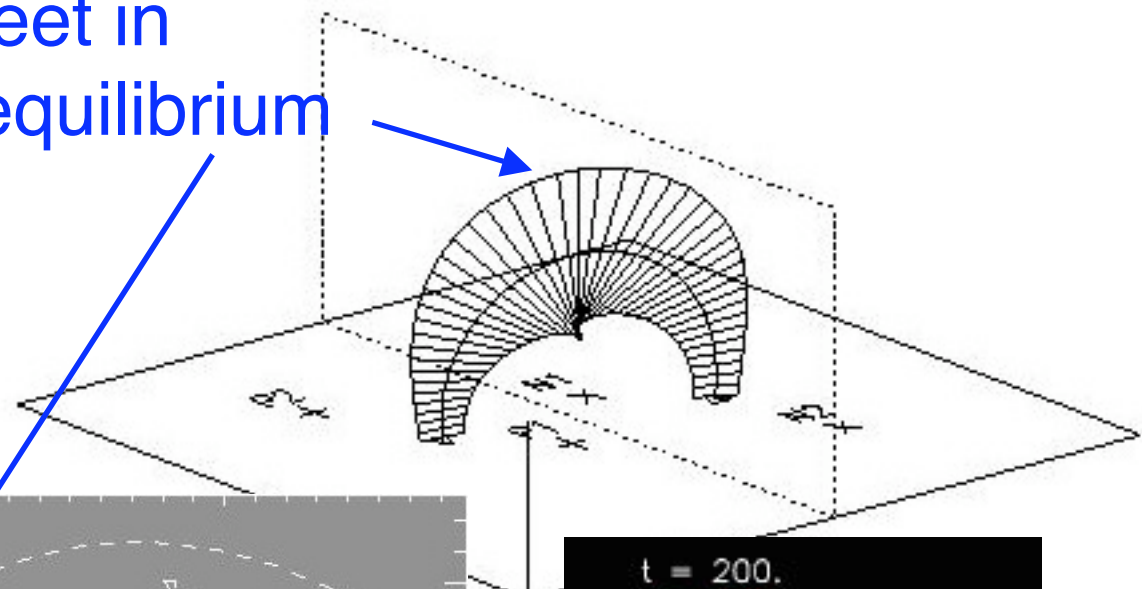


Numerical Test

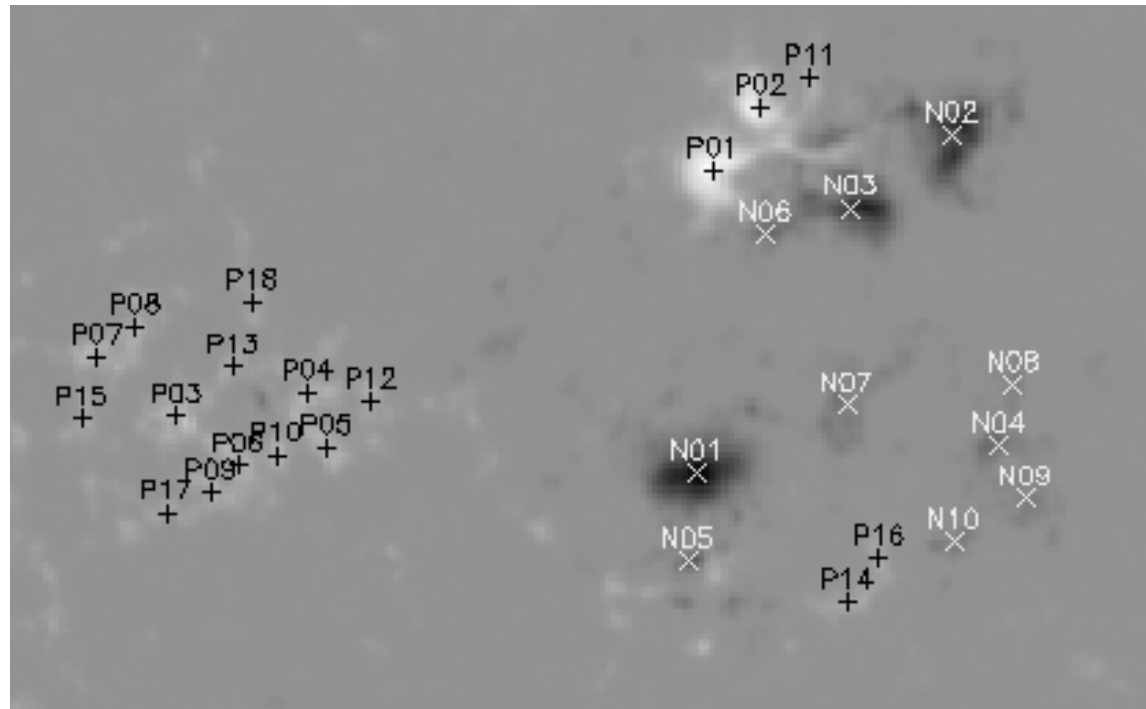
(Longcope & Magara 2003)

Model: Current sheet in flux-constrained equilibrium

Simulation: J/B @ midplane



A complex Example

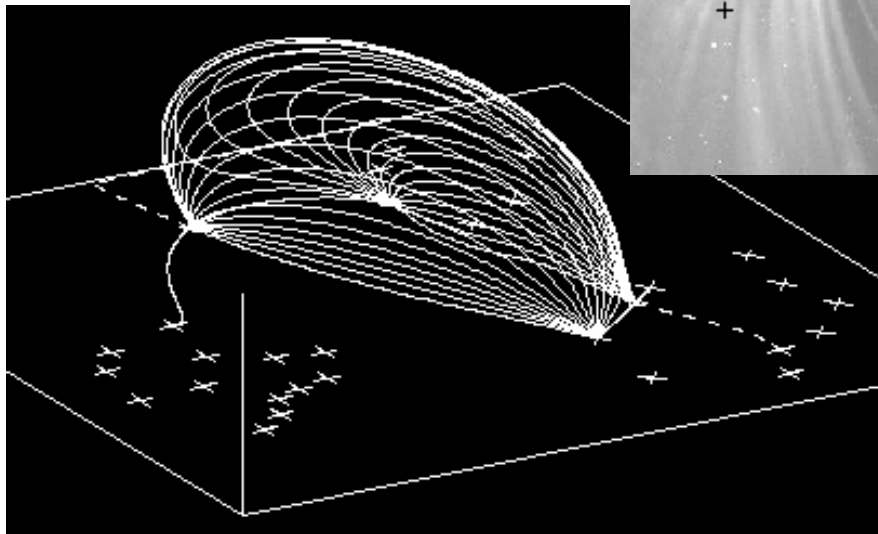
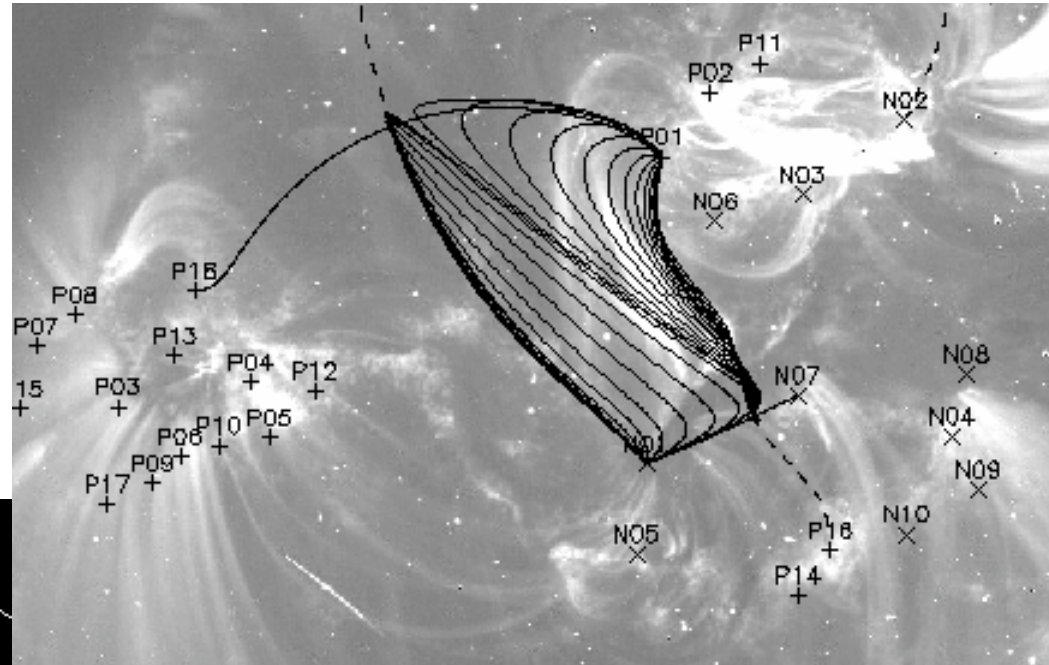


Approximate p-spheric field using discrete sources

The domain of new flux

Emerging bipole
P01-N03

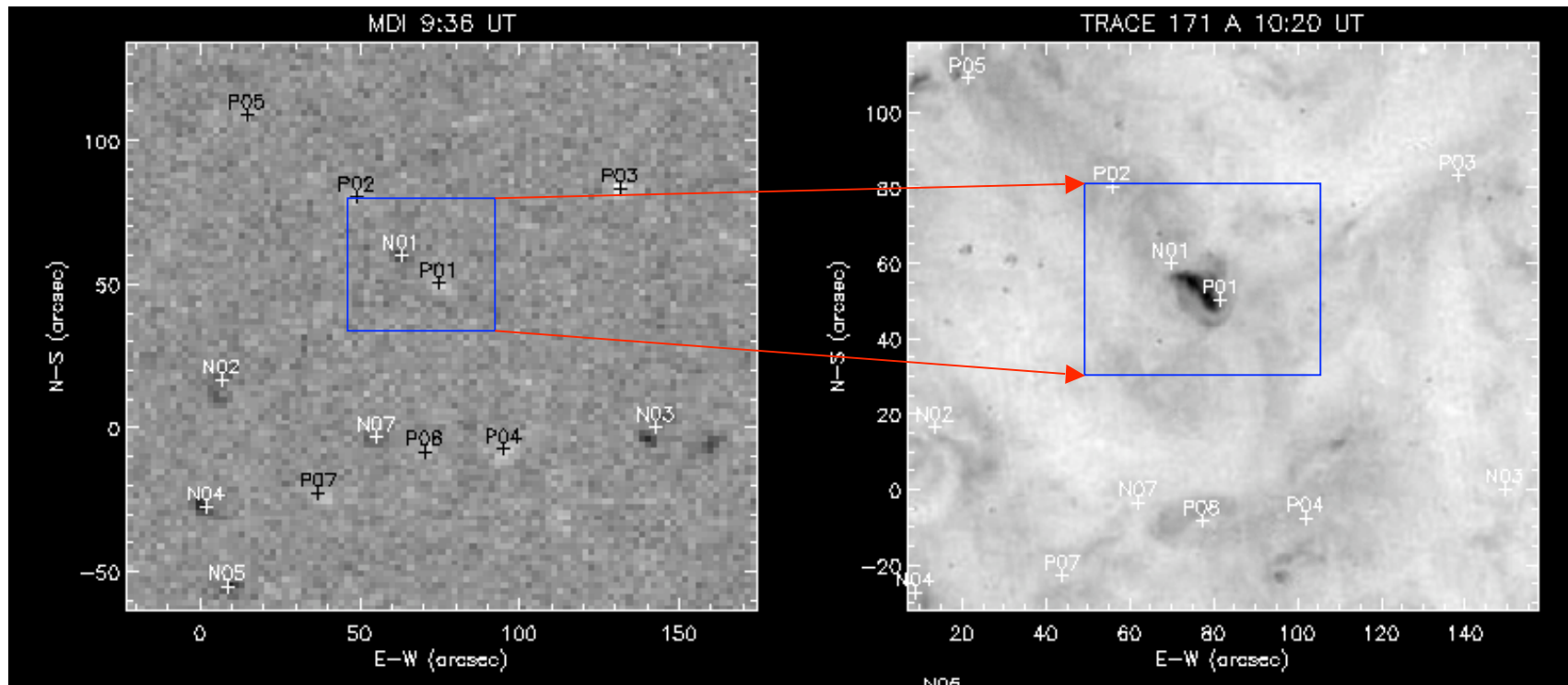
New flux connects
P01-N07



Summary

- 3d reconnection occurs at separators
- Currents accumulate at separators
 - store magnetic energy
- Reconnection there releases energy

A Case Study



(movie)

TRACE & SOI/MDI observations 6/17/98

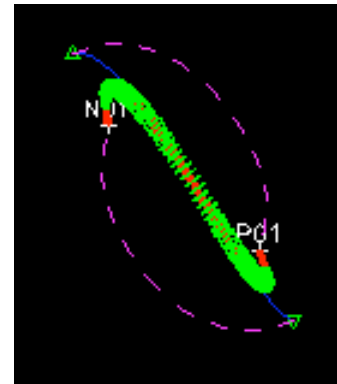
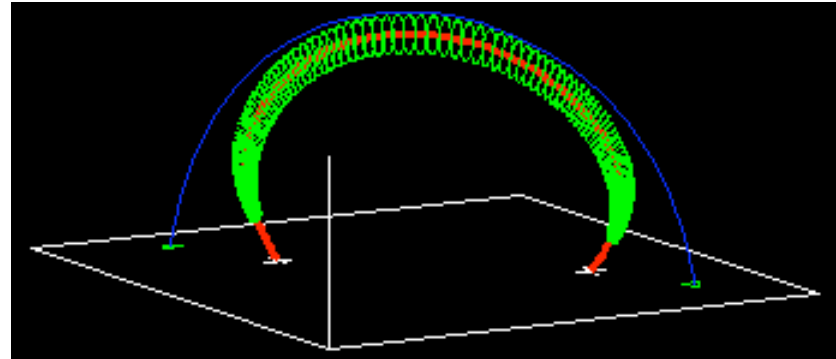
(movie)

(Kankelborg & Longcope 1999)

Post-reconnection Flux Tube

Flux $\Delta\Phi = 1.8 \times 10^{17}$ Mx
Accumulated over

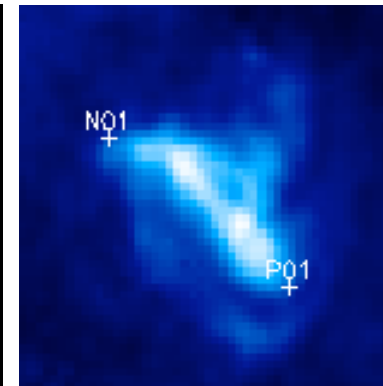
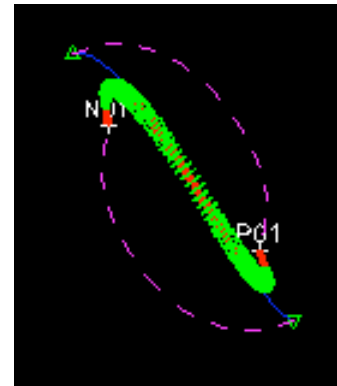
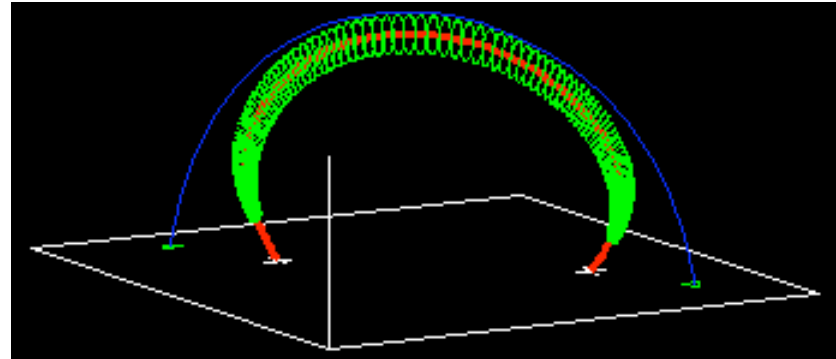
$\Delta t = 20$ min.



Projected
to bipole
location

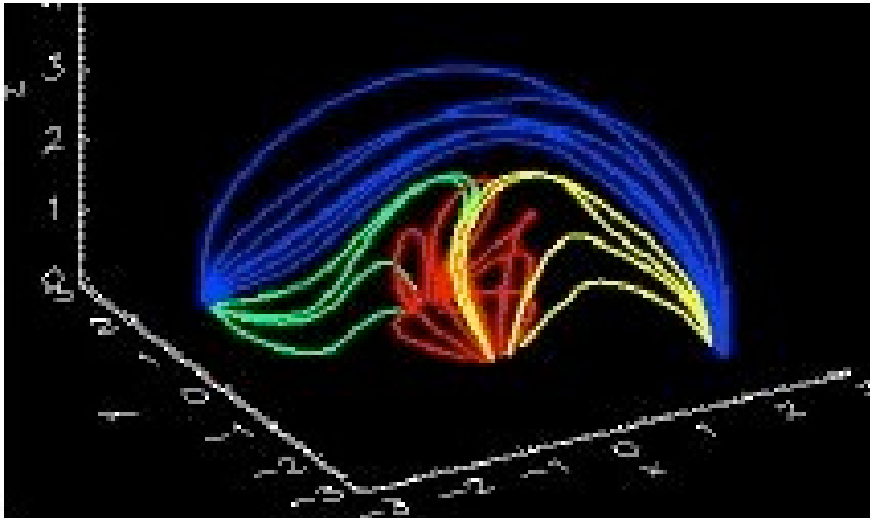
Post-reconnection Flux Tube

Flux $\Delta\Phi = 1.8 \times 10^{17}$ Mx
Accumulated over
 $\Delta t = 20$ min.

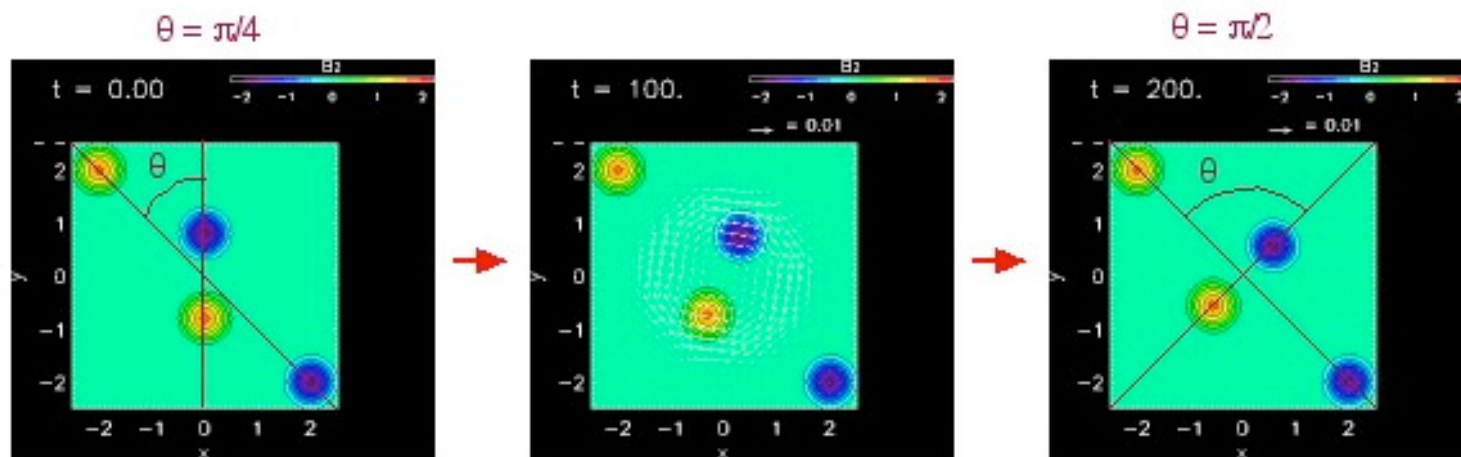


Numerical Test

(Longcope & Magara 2003)



- Initially: potential field
- Move 2 inner sources **slowly**
- Solve 3d MHD eqns. (inside box)



Numerical Test

(Longcope & Magara 2003)

